On the derivation of guaranteed and *p*-robust a posteriori error estimates for the Helmholtz equation

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ANR APOWA kick-off, March 26, 2024

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Model problem

Given $f : \Omega \to \mathbb{C}$, find $u : \Omega \to \mathbb{C}$ such that

$$\begin{cases} -\omega^2 \mu u - \nabla \cdot (\mathbf{A} \nabla u) &= \mu f & \text{in } \Omega, \\ u &= 0 & \text{on } \Gamma_{\mathrm{D}}, \\ \mathbf{A} \nabla u \cdot \mathbf{n} - i \omega \gamma u &= 0 & \text{on } \Gamma_{\mathrm{A}}, \end{cases}$$

where μ , **A**, and γ are given strictly positive (definite) coefficients.



What do we mean by high frequency?

The physical meaning of μ and **A** depends on the application, but the wavespeed is always given by:

$$\mathbf{c_{\min}} := \sqrt{rac{\sigma_{\min}(\mathbf{A})}{\mu}}$$

The (minimal) wavelength is then given by:

$$\lambda := \frac{2\pi}{\omega} c_{\min}.$$

The "important" quantity is $N_{\lambda} := \ell_{\Omega}/\lambda$. High-frequency means that

$$\frac{\omega \ell_{\Omega}}{c_{\min}} \simeq N$$

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 $\mathbf{A} = \mathbf{I}$

 $\Gamma_{\rm A}$

Helmholtz problem and its weak formulation

Recall the Helmholtz problem:

$$\begin{pmatrix} -\omega^2 \mu u - \nabla \cdot (\mathbf{A} \nabla u) &= \mu f & \text{in } \Omega, \\ u &= 0 & \text{on } \Gamma_{\mathrm{D}}, \\ \mathbf{A} \nabla u \cdot \mathbf{n} - i\omega \gamma u &= 0 & \text{on } \Gamma_{\mathrm{A}}. \end{cases}$$

Weak formulation: assuming $f\in L^2(\Omega)$, we seek $u\in H^1_{
m D}(\Omega)$ such that

$$b(\boldsymbol{u},\boldsymbol{v})=(\mu f,\boldsymbol{v}) \quad \forall \boldsymbol{v}\in H^1_{\mathrm{D}}(\Omega),$$

where

$$b(\boldsymbol{u},\boldsymbol{v}) := -\omega^2(\mu\boldsymbol{u},\boldsymbol{v})_{\Omega} - i\omega(\gamma\boldsymbol{u},\boldsymbol{v})_{\Gamma_{\mathrm{A}}} + (\mathbf{A}\boldsymbol{\nabla}\boldsymbol{u},\boldsymbol{\nabla}\boldsymbol{v})_{\Omega}.$$

Helmholtz problem and its weak formulation

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Weak formulation: assuming $f \in L^2(\Omega)$, we seek $u \in H^1_D(\Omega)$ such that

$$b(\mathbf{u},\mathbf{v})=(\mathbf{\mu f},\mathbf{v}) \quad orall \mathbf{v}\in H^1_{\mathrm{D}}(\Omega),$$

where

$$b(\boldsymbol{u},\boldsymbol{v}) := -\omega^2 (\mu \boldsymbol{u},\boldsymbol{v})_{\Omega} - i\omega(\gamma \boldsymbol{u},\boldsymbol{v})_{\Gamma_{\mathrm{A}}} + (\mathbf{A} \nabla \boldsymbol{u}, \nabla \boldsymbol{v})_{\Omega}.$$

Finite element approximation

We consider a mesh \mathcal{T}_h of Ω formed by tetrahedral elements K. Mesh parameter: $h_K := \text{diam } K \leq h$. The coefficients μ, γ, \mathbf{A} are constant inside each element/face.

We introduce the finite element discretization space

$$V_{\boldsymbol{h}} := \left\{ v_{\boldsymbol{h}} \in H^{1}_{\mathrm{D}}(\Omega) \mid v_{\boldsymbol{h}} |_{\mathcal{K}} \in \mathbb{P}_{\boldsymbol{p}}(\mathcal{K}) \; \forall \mathcal{K} \in \mathcal{T}_{\boldsymbol{h}} \right\}$$

with $p \geq 1$.



What do we mean by a fine mesh?

"The mesh is fine" means that

$$N_{
m dofs/\lambda} = \lambda \left/ \frac{h}{p} \simeq \left(\frac{\omega h}{c_{
m min} p} \right)^{-1}$$

is large.



Finite element approximation

Recall that \boldsymbol{u} is the only element of $H^1_{\mathrm{D}}(\Omega)$ such that

$$b(\mathbf{u}, \mathbf{v}) = (\mu \mathbf{f}, \mathbf{v}) \quad \forall \mathbf{v} \in H^1_{\mathrm{D}}(\Omega).$$

Analogously, the finite element solution $u_h \in V_h$ is such that

$$b(\boldsymbol{u}_h, \boldsymbol{v}_h) = (\mu \boldsymbol{f}, \boldsymbol{v}_h) \quad \forall \boldsymbol{v}_h \in \boldsymbol{V}_h. \tag{1}$$

In this talk, we are interested in estimating the error

 $e_h := u - u_h$



A priori error estimates

A priori estimate

Assume that $\omega h/c_{\min}p \leq \mathscr{C}_1$. Then

$$\| \boldsymbol{\nabla} \boldsymbol{e}_{h} \|_{\boldsymbol{A},\Omega} \leq \mathscr{C}_{2} \left(\frac{\omega h}{c_{\min} p} \right)^{p}$$



F. Ihlenburg and I. Babuška, SIAM J. Numer. Anal., 1997.

J.M. Melenk and S.A. Sauter, *Math. Comp.*, 2010.

T. Chaumont-Frelet and S. Nicaise, IMA J. Numer. Anal., 2019.

Some limitations:

- The above result requires important regularity assumptions.
- The error estimate is not always applicable.
- The constant \mathscr{C}_2 is *not computable* in general.

A priori error estimates

A priori estimates provide qualitative upper bounds.

They are important as they show that the method converges. They also indicate how fast the convergence happens.

They are not suited to *quantitatively* estimate the error in practice.

A posteriori estimates

(Ideal) a posteriori estimate	
	$\ oldsymbol{ abla} e_h \ _{oldsymbol{A},\Omega} \leq \eta.$

Here η is a fully-computable real number called an "error estimator". This quantity is computed as a *local* post-processing of u_h , i.e. $\eta = \eta(u_h)$. There are *no generic constants*. We have a *guaranteed* error estimate.



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The low-frequency case

We first consider the low frequency limit where $\omega = 0$.

The problem then reads: find ${\it u}:\Omega\to \mathbb{C}$ such that

$$\begin{pmatrix} -\boldsymbol{\nabla} \cdot (\boldsymbol{A} \boldsymbol{\nabla} \boldsymbol{u}) &= \boldsymbol{\mu} \boldsymbol{f} & \text{in } \Omega, \\ \boldsymbol{u} &= \boldsymbol{0} & \text{on } \boldsymbol{\Gamma}_{\mathrm{D}}, \\ \boldsymbol{A} \boldsymbol{\nabla} \boldsymbol{u} \cdot \boldsymbol{n} &= \boldsymbol{0} & \text{on } \boldsymbol{\Gamma}_{\mathrm{A}}. \end{cases}$$

For the sake of simplicity, let $f = f_h \in \mathbb{P}_p(\mathcal{T}_h)$.



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At the continuous level, the ideal flux $\boldsymbol{\sigma} := -\boldsymbol{\mathsf{A}} \boldsymbol{\nabla} \boldsymbol{\mathsf{u}}$ satisfies

$$\sigma = \arg \min_{\substack{\boldsymbol{\tau} \in \boldsymbol{H}_{\Gamma_{A}}(\operatorname{div},\Omega) \\ \boldsymbol{\nabla} \cdot \boldsymbol{\tau} = \mu f_{h} \text{ in } \Omega}} \| \mathbf{A}^{-1} \boldsymbol{\tau} + \boldsymbol{\nabla} \boldsymbol{u} \|_{\mathbf{A},\Omega}.$$

The best-possible computable flux directly mimics this definition at the discrete level

$$\sigma_h := \arg \min_{\substack{\boldsymbol{\tau}_h \in \boldsymbol{W}_h \\ \boldsymbol{\nabla} \cdot \boldsymbol{\tau}_h = \mu f_h \text{ in } \Omega}} \| \mathbf{A}^{-1} \boldsymbol{\tau}_h + \boldsymbol{\nabla} \boldsymbol{u}_h \|_{\mathbf{A}, \Omega}.$$

Consider the set of "hat functions" $\{\psi^a\}_{a \in \mathcal{V}_h}$ of the mesh. We then have

$$\sum_{\boldsymbol{a}\in\mathcal{V}_h}\psi^{\boldsymbol{a}}=1.$$



The ideal flux $\boldsymbol{\sigma} := -\mathbf{A} \nabla u$ can be decomposed as

$$\boldsymbol{\sigma} = \sum_{\boldsymbol{a}\in\mathcal{V}_h} \boldsymbol{\sigma}^{\boldsymbol{a}}, \qquad \boldsymbol{\sigma}^{\boldsymbol{a}} = -\psi^{\boldsymbol{a}} \mathbf{A} \nabla u.$$

Easy computations show that

$$\sigma^a \cdot \mathbf{n} = 0$$
 on $\partial \omega^a$, $\nabla \cdot \sigma^a = \psi^a \mu f_h - \mathbf{A} \nabla u \cdot \nabla \psi^a$ in ω^a .

Consider the set of "hat functions" $\{\psi^a\}_{a \in \mathcal{V}_h}$ of the mesh. We then have

$$\sum_{a \in \mathcal{V}_b} \psi^a = 1.$$



The ideal flux $\boldsymbol{\sigma} := -\mathbf{A} \nabla \boldsymbol{u}$ can be decomposed as

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Easy computations show that

$$\sigma^a \cdot \mathbf{n} = 0$$
 on $\partial \omega^a$, $\nabla \cdot \sigma^a = \psi^a \mu f_b - \mathbf{A} \nabla u \cdot \nabla \psi^a$ in ω^a .

We have shown that

$$\sigma = \sum_{a \in \mathcal{V}_h} \sigma^a$$

and

$$\begin{aligned} \boldsymbol{\sigma}^{\boldsymbol{a}} &= \arg \min_{\substack{\boldsymbol{\tau} \in \boldsymbol{H}_0(\operatorname{div}, \omega^{\boldsymbol{a}}) \\ \boldsymbol{\nabla} \cdot \boldsymbol{\tau} = \psi^{\boldsymbol{a}} \mu f_h - \boldsymbol{A} \boldsymbol{\nabla} u \cdot \boldsymbol{\nabla} \psi^{\boldsymbol{a}} \text{ in } \omega^{\boldsymbol{a}} } \| \boldsymbol{A}^{-1} \boldsymbol{\tau} + \psi^{\boldsymbol{a}} \boldsymbol{\nabla} u \|_{\boldsymbol{A}, \omega^{\boldsymbol{a}}}. \end{aligned}$$

Implicit characterization that we can mimic at the discrete level!

We have shown that

$$\boldsymbol{\sigma} = \sum_{\boldsymbol{a} \in \mathcal{V}_h} \boldsymbol{\sigma}^{\boldsymbol{a}}$$

and

$$\begin{aligned} \boldsymbol{\sigma}^{a} &= \arg \min_{\substack{\boldsymbol{\tau} \in \boldsymbol{H}_{0}(\operatorname{div}, \omega^{a}) \\ \boldsymbol{\nabla} \cdot \boldsymbol{\tau} = \psi^{a} \mu f_{h} - \mathbf{A} \boldsymbol{\nabla} u \cdot \boldsymbol{\nabla} \psi^{a} \text{ in } \omega^{a} } \| \mathbf{A}^{-1} \boldsymbol{\tau} + \psi^{a} \boldsymbol{\nabla} u \|_{\mathbf{A}, \omega^{a}}. \end{aligned}$$

Implicit characterization that we can mimic at the discrete level!

We thus set

$$\boldsymbol{\sigma}_h := \sum_{\boldsymbol{a} \in \mathcal{V}_h} \boldsymbol{\sigma}_h^{\boldsymbol{a}}$$

with

$$\begin{split} \boldsymbol{\sigma}_h^{\boldsymbol{a}} := \arg \min_{ \substack{\boldsymbol{\tau}_h \in \boldsymbol{H}_0(\operatorname{div}, \omega^a) \cap \boldsymbol{W}_h \\ \boldsymbol{\nabla} \cdot \boldsymbol{\tau}_h = \psi^a \mu f_h - \mathbf{A} \boldsymbol{\nabla} \boldsymbol{u}_h \cdot \boldsymbol{\nabla} \psi^a \text{ in } \omega^a } \| \mathbf{A}^{-1} \boldsymbol{\tau}_h + \psi^a \boldsymbol{\nabla} \boldsymbol{u}_h \|_{\mathbf{A}, \omega^a} . \end{split}$$

The compatibility condition

$$(\psi^a \mu f_h - \mathbf{A} \nabla u_h \cdot \nabla \psi^a, 1)_{\omega^a} = (\mu f_h, \psi^a)_{\Omega} - (\mathbf{A} \nabla u_h, \nabla \psi^a)_{\Omega} = 0$$

holds true since u_h is the discrete solution and $\psi^a \in V_h$.

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$$\boldsymbol{\sigma}_h := \sum_{\boldsymbol{a} \in \mathcal{V}_h} \boldsymbol{\sigma}_h^{\boldsymbol{a}}$$

with

$$\begin{aligned} \boldsymbol{\sigma}_h^a := \arg \min_{ \substack{\boldsymbol{\tau}_h \in \boldsymbol{H}_0(\operatorname{div}, \omega^a) \cap \boldsymbol{W}_h \\ \boldsymbol{\nabla} \cdot \boldsymbol{\tau}_h = \psi^a \mu f_h - \mathbf{A} \boldsymbol{\nabla} \boldsymbol{u}_h \cdot \boldsymbol{\nabla} \psi^a \text{ in } \omega^a } \| \mathbf{A}^{-1} \boldsymbol{\tau}_h + \psi^a \boldsymbol{\nabla} \boldsymbol{u}_h \|_{\mathbf{A}, \omega^a}. \end{aligned}$$

The compatibility condition

$$(\psi^{a}\mu f_{h} - \mathbf{A}\nabla u_{h} \cdot \nabla \psi^{a}, 1)_{\omega^{a}} = (\mu f_{h}, \psi^{a})_{\Omega} - (\mathbf{A}\nabla u_{h}, \nabla \psi^{a})_{\Omega} = 0$$

holds true since u_h is the discrete solution and $\psi^a \in V_h$.

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Summary of the localization process

Step 1: solve a set of small, uncoupled linear systems

$$\begin{split} \boldsymbol{\sigma}_h^{\boldsymbol{a}} &:= \arg \min_{ \substack{\boldsymbol{\tau}_h \in \boldsymbol{H}_0(\operatorname{div}, \omega^a) \cap \boldsymbol{W}_h \\ \boldsymbol{\nabla} \cdot \boldsymbol{\tau}_h = \psi^a \mu f_h - \boldsymbol{A} \boldsymbol{\nabla} \boldsymbol{u}_h \cdot \boldsymbol{\nabla} \psi^a \text{ in } \omega^a } } \| \boldsymbol{A}^{-1} \boldsymbol{\tau}_h + \psi^a \boldsymbol{\nabla} \boldsymbol{u}_h \|_{\boldsymbol{A}, \omega^a}. \end{split}$$

Step 2: assemble these local contributions

$$\boldsymbol{\sigma}_h := \sum_{\boldsymbol{a} \in \mathcal{V}_h} \boldsymbol{\sigma}_h^{\boldsymbol{a}}.$$

Step 3: compute the estimator

$$\eta := \|\mathbf{A}^{-1}\boldsymbol{\sigma}_h + \boldsymbol{\nabla}\boldsymbol{u}_h\|_{\mathbf{A},\Omega}.$$

Step 4: enjoy the guaranteed estimate by the Prager-Synge (in)equality

$$\|\nabla e_h\|_{\mathbf{A},\Omega} \leq \eta.$$

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Efficiency

We can show that

$$\eta \leq C_{ ext{eff}} \| \boldsymbol{
abla} \boldsymbol{e}_h \|_{\mathbf{A},\Omega},$$

where C_{eff} only depends on:

- the "flatness" of the tetrahedra in the mesh,
- the "contrasts" in the coefficients.

An important aspect is that C_{eff} does not depend on the polynomial degree p. Local efficiency can be shown as well.

- P. Braess, V. Pillwein, and J. Schöberl, *CMAME*, 2009.
- A. Ern and M. Vohralík, SINUM, 2015 & Math. Comp., 2020.

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The high-frequency case

Back to our original problem

$$\begin{cases} -\omega^2 \mu u - \nabla \cdot (\mathbf{A} \nabla u) &= \mu f & \text{in } \Omega, \\ u &= 0 & \text{on } \Gamma_D, \\ \mathbf{A} \nabla u \cdot \mathbf{n} - i \omega \gamma u &= 0 & \text{on } \Gamma_A. \end{cases}$$

Recall

$$b(\boldsymbol{u},\boldsymbol{v}) = -\omega^2(\mu\boldsymbol{u},\boldsymbol{v})_{\Omega} - i\omega(\gamma\boldsymbol{u},\boldsymbol{v})_{\Gamma_{\mathrm{A}}} + (\mathbf{A}\boldsymbol{\nabla}\boldsymbol{u},\boldsymbol{\nabla}\boldsymbol{v})_{\Omega}.$$

Energy norm and lack of coercivity

We will consider the "balanced" norm

$$|\!|\!| \boldsymbol{v} |\!|\!|_{\boldsymbol{\omega},\Omega}^2 := \boldsymbol{\omega}^2 |\!| \boldsymbol{v} |\!|_{\boldsymbol{\mu},\Omega}^2 + |\!| \boldsymbol{\nabla} \boldsymbol{v} |\!|_{\boldsymbol{A},\Omega}^2.$$

The sesquilinear form *b* is *not coercive*.

Instead we have the "Gårding" inequality

$$\operatorname{\mathsf{Re}} b(v,v) = \|\nabla v\|_{\mathbf{A},\Omega}^2 - \omega^2 \|v\|_{\mu,\Omega}^2 = \|v\|_{\omega,\Omega}^2 - 2\omega^2 \|v\|_{\mu,\Omega}^2.$$

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Letting $\boldsymbol{\sigma} := -\mathbf{A} \nabla \boldsymbol{u}$, we have

$$\boldsymbol{\sigma} \cdot \boldsymbol{n} = -i\omega\gamma \boldsymbol{u}$$
 on $\Gamma_{\rm A}$ $\boldsymbol{\nabla} \cdot \boldsymbol{\sigma} = \mu f_{\boldsymbol{h}} + \omega^2 \mu \boldsymbol{u}$ in Ω .

Hence, natural requirements for σ_h are

$$\boldsymbol{\sigma}_h \cdot \boldsymbol{n} = -i\omega\gamma \boldsymbol{u}_h$$
 on Γ_A $\boldsymbol{\nabla} \cdot \boldsymbol{\sigma}_h = \mu f_h + \omega^2 \mu \boldsymbol{u}_h$ in Ω .

The "low frequency" reconstruction algorithm directly extends.

Prager–Synge inequality

Let $v \in H^1_{\Gamma_D}(\Omega)$. We have $\begin{aligned} b(\boldsymbol{e}_h, \boldsymbol{v}) &= (\mu f_h, \boldsymbol{v})_\Omega - b(\boldsymbol{u}_h, \boldsymbol{v}) \\ &= (\mu f_h + \omega^2 \mu \boldsymbol{u}_h, \boldsymbol{v})_\Omega + i\omega(\gamma \boldsymbol{u}_h, \boldsymbol{v})_{\Gamma_A} - (\mathbf{A} \nabla \boldsymbol{u}_h, \nabla \boldsymbol{v})_\Omega \\ &= (\nabla \cdot \boldsymbol{\sigma}_h, \boldsymbol{v})_\Omega - (\boldsymbol{\sigma}_h \cdot \boldsymbol{n}, \boldsymbol{v})_{\Gamma_A} - (\mathbf{A} \nabla \boldsymbol{u}_h, \nabla \boldsymbol{v})_\Omega \\ &= -(\boldsymbol{\sigma}_h + \mathbf{A} \nabla \boldsymbol{u}_h, \nabla \boldsymbol{v})_\Omega. \end{aligned}$

Prager–Synge inequality

$$\|m{b}(m{e}_{m{h}},m{v})\|\leq m{\eta}\|m{
abla}m{v}\|_{m{A},\Omega} \hspace{1em}orall\,v\in H^1_{\Gamma_{
m D}}(\Omega)$$

So far, so good!

What's the matter?

Here, we do not have

$$\| oldsymbol{
abla} e_h \|_{oldsymbol{A},\Omega}^2 \leq |b(e_h,e_h)|,$$

which is a major issue!

Instead, we only have the "Gårding" inequality

$$\operatorname{\mathsf{Re}} b(\boldsymbol{e}_h, \boldsymbol{e}_h) \geq \| \boldsymbol{e}_h \|_{\omega,\Omega}^2 - 2\omega^2 \| \boldsymbol{e}_h \|_{\mu,\Omega}^2.$$

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Stability constant

For $g \in L^2(\Omega)$, let \mathscr{S}^*g denote the unique element of $H^1_{\Gamma_{\Omega}}(\Omega)$ such that

$$b(w, \mathscr{S}^{\star}g) = 2\omega^2(\mu w, g)_{\Omega} \quad \forall w \in H^1_{\Gamma_{\mathrm{D}}}(\Omega)$$

and let

$$\mathscr{C}_{\mathrm{st}} := rac{1}{\omega} \max_{\substack{ g \in L^2(\Omega) \\ \|g\|_{\mu,\Omega} = 1}} \| oldsymbol{
abla}(\mathscr{S}^\star g) \|_{\mathbf{A},\Omega}.$$

 \mathscr{C}_{st} is the best constant such that

$$\| oldsymbol{
abla}(\mathscr{S}^{\star}g) \|_{oldsymbol{A},\Omega} \leq \mathscr{C}_{\mathrm{st}} \omega \| g \|_{\mu,\Omega} \quad orall g \in L^2(\Omega).$$

It is closely related to resolvant estimates.

Making up for the lack of coercivity

By definition, we have

$$b(w, \mathscr{S}^{\star} \boldsymbol{e}_{h}) = 2\omega^{2}(\mu w, \boldsymbol{e}_{h}) \quad \forall w \in H^{1}_{\Gamma_{\mathrm{D}}}(\Omega).$$

Hence, in particular,

$$b(\boldsymbol{e}_h, \mathscr{S}^{\star} \boldsymbol{e}_h) = 2\omega^2 \|\boldsymbol{e}_h\|_{\mu,\Omega}^2,$$

which is exactly the "bad" term the Gårding inequality:

$$\operatorname{Re} b(\boldsymbol{e}_h, \boldsymbol{e}_h) = \|\boldsymbol{e}_h\|_{\omega,\Omega}^2 - 2\omega^2 \|\boldsymbol{e}_h\|_{\mu,\Omega}^2.$$
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Making up for the lack of coercivity

Using Prager-Synge inequality, we have

$$\|oldsymbol{e}_h\|_{\omega,\Omega}^2 = \mathsf{Re}\,b(oldsymbol{e}_h,oldsymbol{e}_h+\mathscr{S}^{\star}oldsymbol{e}_h) \leq \eta \|oldsymbol{
abla}(oldsymbol{e}_h+\mathscr{S}^{\star}oldsymbol{e}_h)\|_{\mathbf{A},\Omega}$$

It follows that

$$\begin{split} \left\| \left| e_h \right| \right|_{\omega,\Omega}^2 &\leq \eta \left(\left\| \nabla e_h \right\|_{\mathbf{A},\Omega} + \left\| \nabla (\mathscr{S}^* e_h) \right\|_{\mathbf{A},\Omega} \right) \\ &\leq \eta \left(\left\| \nabla e_h \right\|_{\mathbf{A},\Omega} + \mathscr{C}_{\mathrm{st}} \omega \left\| e_h \right\|_{\mu,\Omega} \right) \\ &\leq \eta \max(1,\mathscr{C}_{\mathrm{st}}) \left\| \left| e_h \right\|_{\omega,\Omega}, \end{split}$$

and

$$\| e_h \|_{\omega,\Omega} \leq \max(1, \mathscr{C}_{\mathrm{st}}) \eta.$$



Coarse error estimate

We obtained the following error estimate:

Coarse error estimate $\| \boldsymbol{e}_{\boldsymbol{h}} \| \leq \max(1, \mathscr{C}_{\mathrm{st}}) \boldsymbol{\eta}$

\mathscr{C}_{st} is the best constant such that:

Stability constant $\forall \mathbf{g} \in L^2(\Omega).$ $\|\nabla(\mathscr{S}^{\star}g)\|_{\mathbf{A},\Omega} \leq \mathscr{C}_{\mathrm{st}}\omega\|g\|_{\mu,\Omega}$

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Efficiency

We can show that

$$\eta \leq C_{\text{eff}} \left(1 + \max_{K \in \mathcal{T}_h} \frac{\omega h_K}{c_{\min K} \rho} \right) \left\| e_h \right\|_{\omega,\Omega},$$

where $c_{\min K}$ is the wavespeed in the element K. For any reasonable discretization, we have

$$\frac{\omega h_K}{c_{\min K} p} \le 1,$$

so that in practice

$$\eta \leq C_{\mathrm{eff}} \left\| e_h \right\|_{\omega,\Omega},$$

where $C_{\rm eff}$ only depends on the elements "flatness" and the contrasts.

W. Dörfler, S. Sauter, Comput. Meth. Appl. Math., 2013.

T. Chaumont-Frelet, A. Ern, and M. Vohralík, Numer. Math., 2021.

The problem with the coarse error estimate

Recall that

$$\eta \leq C_{\mathrm{eff}} \left\| \left\| oldsymbol{e}_h
ight\|_{\omega,\Omega}
ight| = \max(1,\mathscr{C}_{\mathrm{st}})\eta.$$

We have $\mathit{C}_{
m eff}\simeq 1$ and $\mathscr{C}_{
m st}\gtrsim \omega\ell_\Omega/\mathit{c}_{
m min}$, so that

$$\eta \lesssim \| \| \boldsymbol{e}_h \| \|_{\omega,\Omega} \lesssim \frac{\omega \ell_\Omega}{c_{\min}} \eta.$$

In practice, the coarse error estimate will largely overestimate the error in the high-frequency regime.

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The approximation factor

We introduce
$$\mathscr{C}_{\mathrm{ap}} := \frac{1}{\omega} \max_{\substack{g \in L^2(\Omega) \\ \|g\|_{\mu,\Omega} = 1}} \min_{v_h \in V_h} \|\nabla(\mathscr{S}^*g - v_h)\|_{\mathbf{A},\Omega}.$$

Approximability

For all $g \in L^2(\Omega)$, there exists $v_h^{\star} \in V_h$ such that

$$\| oldsymbol{\nabla} (\mathscr{S}^{\star} g - v^{\star}_{h}) \|_{oldsymbol{A},\Omega} \leq \mathscr{C}_{\mathrm{ap}} \omega \| g \|_{\mu,\Omega}.$$

Using Galerkin orthogonality

Recall that

$$|\!|\!|\!|_{\boldsymbol{\omega},\Omega}^2 = \operatorname{Re} b(\underline{e}_h, \underline{e}_h + \mathscr{S}^* \underline{e}_h).$$

By Galerkin orthogonality, we have

$$\begin{split} \|\boldsymbol{e}_{h}\|^{2}_{\boldsymbol{\omega},\Omega} &= \operatorname{\mathsf{Re}} b(\boldsymbol{e}_{h},\boldsymbol{e}_{h}) + \operatorname{\mathsf{Re}} b(\boldsymbol{e}_{h},\mathscr{S}^{\star}\boldsymbol{e}_{h}) \\ &= \operatorname{\mathsf{Re}} b(\boldsymbol{e}_{h},\boldsymbol{e}_{h}) + \operatorname{\mathsf{Re}} b(\boldsymbol{e}_{h},\mathscr{S}^{\star}\boldsymbol{e}_{h} - \boldsymbol{v}_{h}^{\star}) \\ &\leq \eta \left(\|\boldsymbol{\nabla}\boldsymbol{e}_{h}\|_{\mathbf{A},\Omega} + \|\boldsymbol{\nabla}(\mathscr{S}^{\star}\boldsymbol{e}_{h} - \boldsymbol{v}_{h}^{\star})\|_{\mathbf{A},\Omega} \right) \\ &\leq \eta \max(1,\mathscr{C}_{\operatorname{ap}}) \|\|\boldsymbol{e}_{h}\|_{\boldsymbol{\omega},\Omega} \,. \end{split}$$



Sharp error estimate

Sharp error estimate

$$\|\| \boldsymbol{e}_{\boldsymbol{h}} \|\|_{\omega,\Omega} \leq \max(1, \mathscr{C}_{\mathrm{ap}}) \eta.$$

Approximation factor

$$\mathscr{C}_{\mathrm{ap}} := rac{1}{\omega} \max_{\substack{g \in L^2(\Omega) \ \|g\|_{\mu,\Omega} = 1}} \min_{\mathbf{v}_h \in V_h} \| \mathbf{\nabla} (\mathscr{S}^{\star}g - \mathbf{v}_h) \|_{\mathbf{A},\Omega} o 0.$$

W. Dörfler, S. Sauter, Comput. Meth. Appl. Math., 2013.

T. Chaumont-Frelet, A. Ern, and M. Vohralík, Numer. Math., 2021.

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Takeaways

The "equilibration" technology is the same as for low frequencies.

Coarse error estimate

$$\|m{e}_{m{h}}\|_{\omega,\Omega} \leq \max(1,\mathscr{C}_{\mathrm{st}})\eta \quad \mathscr{C}_{\mathrm{st}} \gtrsim \omega \ell_\Omega/c_{\mathsf{min}}$$

Sharp error estimate

$$\|\!|\!|\!|\!|_{\omega,\Omega} \leq \max(1,\mathscr{C}_{\mathrm{ap}})\eta \quad \mathscr{C}_{\mathrm{ap}} o 0$$

Efficiency

$$\eta \leq C_{ ext{eff}} \left(1 + rac{\omega h}{c_{ ext{min}}
ho}
ight) \|\!|\!| e_h |\!|\!|_{\omega,\Omega}$$

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The stability constant

The stability constant is defined by

$$\mathscr{C}_{\mathrm{st}} := rac{1}{\omega} \max_{\substack{g \in L^2(\Omega) \ \|g\|_{\mu,\Omega} = 1}} \|oldsymbol{
abla}(\mathscr{S}^{\star}g)\|_{\mathbf{A},\Omega}.$$

It is only related to the PDE and independent of the numerical scheme.

The stability constant $\mathscr{C}_{\mathrm{st}}$ The approximation factor $\mathscr{C}_{\mathrm{ap}}$

Qualitative behaviour

It is known that we have "at least":

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$$\mathscr{C}_{\mathrm{st}}\gtrsim rac{\omega\ell_\Omega}{c_{\mathsf{min}}}.$$

 $\mathscr{C}_{\mathrm{st}} \lesssim \frac{\omega \ell_{\Omega}}{c}.$

For non-trapping settings (the "easier" scenario), we have

If strong traping happens, "extreme" behaviors can occur

$$\mathscr{C}_{\mathrm{st}}\gtrsim\exp\left(lpharac{\omega\ell_{\Omega}}{c_{\min}}
ight)$$

for "some" frequencies. For "most frequencies"

$$\mathscr{C}_{\mathrm{st}}\gtrsim \left(rac{\omega\ell_\Omega}{c_{\mathsf{min}}}
ight)^eta.$$

D. Lafontaine, E.A. Spence, and J. Wunsch, Comm. Pure Appl. Math. 2021

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m st}$ The approximation factor $\mathscr{C}_{
m ap}$

Quantitative estimate for star-shaped non-trapping obstacles

 $\Omega := (-\ell/2, \ell/2)^3$ is a cube centered at the origin. $D \subset \Omega$ is star shaped with respect to the origin.

Assume that $\gamma = 1$ and that $\mu = 1$ and $\mathbf{A} = \mathbf{I}$ in $\Omega \setminus D$.

Assume that $\mu = \mu_D \ge 1$ and $\mathbf{A} = \mathbf{A}_D \preceq \mathbf{I}$ in D.

This describes an obstacle made of a material with a "slow" wavespeed.





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Quantitative estimate for star-shaped non-trapping obstacles

Guaranteed upper bound

$$\mathscr{C}_{\mathrm{st}} \leq 6 + rac{3+\sqrt{3}}{\sqrt{3}} rac{\omega \ell_\Omega}{c_{\mathsf{min}}}$$

The proof relies on a "Morawetz multiplier": multiply the PDE by $x \cdot \nabla u$ and integrate by parts until it works!



J.M. Melenk, PhD thesis, 1995.

- H. Barucq, T. Chaumont-Frelet, and C. Gout, Math. Comp., 2016.
- T. Chaumont-Frelet, A. Ern, and M. Vohralík, *Numer. Math.* 2021.

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The approximation factor

The approximation factor is defined by

$$\mathscr{C}_{\mathrm{ap}} := rac{1}{\omega} \max_{\substack{g \in L^2(\Omega) \ \|g\|_{\mu,\Omega} = 1}} \min_{\mathsf{v}_h \in \mathsf{V}_h} \| \mathbf{\nabla} (\mathscr{S}^\star g - \mathsf{v}_h) \|_{\mathbf{A},\Omega}.$$

It depends on both the PDE and the approximation space V_h .

Assuming that $\mathbf{A} = \mathbf{I}$, Ω is convex, and \mathscr{C}_{st} is known, we can control it.

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Idea one: explicit interpolation error

R. Arcangeli and J.L. Gout, RAIRO Numer. Anal., 1976. If $v \in H^2(\Omega)$, let $I_h^1 v \in V_h$ denotes its first-order Lagrange interpolant:

$$\|oldsymbol{
abla}(oldsymbol{
u}-oldsymbol{I}_h^1oldsymbol{
u})\|_{oldsymbol{A},\Omega}\leq \mathscr{C}_{\mathcal{T},\mathrm{i}}h\|oldsymbol{
abla}^2oldsymbol{
u}\|_{\Omega},$$

with a constant $\mathscr{C}_{\mathcal{T},i}$ that is easily computable. We then have

$$egin{aligned} &\mathcal{C}_{\mathrm{ap}} := rac{1}{\omega} \max_{\substack{oldsymbol{g} \in L^2(\Omega) \ \|oldsymbol{g}\|_{\mu,\Omega} = 1}} \min_{oldsymbol{v}_h^\star \in V_h} \|oldsymbol{
abla}(\mathscr{S}^\star g - oldsymbol{v}_h^\star)\|_{\mathbf{A},\Omega} \ &\leq rac{1}{\omega} \max_{\substack{oldsymbol{g} \in L^2(\Omega) \ \|oldsymbol{g}\|_{\mu,\Omega} = 1}} \|oldsymbol{
abla}(\mathscr{S}^\star g - I_h^1(\mathscr{S}^\star g)\|_{\mathbf{A},\Omega} \ &\leq rac{1}{\omega} \mathscr{C}_{\mathcal{T},\mathbf{i}} rac{h}{\|oldsymbol{g}\|_{\mu,\Omega} = 1} \|oldsymbol{
abla}^2(\mathscr{S}^\star g)\|_{\Omega}. \end{aligned}$$

Idea two: estimation of the Hessian norm

P. Grisvard, 1985.

T. Chaumont-Frelet, S. Nicaise, and J. Tomezyk, *Comm. Pure Appl. Anal.*, 2020. Because Ω is convex and $\gamma = 1$, we have

$$\|oldsymbol{
abla}^2(\mathscr{S}^{\star} {oldsymbol{g}})\|_{\Omega} \leq \|\Delta(\mathscr{S}^{\star} {oldsymbol{g}})\|_{\Omega}.$$

Then, we use the facts that

$$-\Delta(\mathscr{S}^{\star}g) = 2\mu\omega^{2}g + \mu\omega^{2}\mathscr{S}^{\star}g$$

and

$$\omega \| \mathscr{S}^{\star} g \|_{\mu,\Omega} \leq 2 \mathscr{C}_{\mathrm{st}} \| g \|_{\mu,\Omega}$$

to show that

$$\| \nabla^2 (\mathscr{S}^{\star} g) \|_{\Omega} \leq 2 rac{\omega}{c_{\min}} \left(1 + \mathscr{C}_{\mathrm{st}}
ight) \| g \|_{\mu,\Omega}.$$

Explicit control of the approximation factor

Recall that

$$\mathscr{C}_{\mathrm{ap}} \leq rac{1}{\omega} \mathscr{C}_{\mathcal{T}, \mathrm{i}} rac{h}{\|g\|_{\mu, \Omega} = 1} \|oldsymbol{
abla}^2(\mathscr{S}^{\star}g)\|_{\Omega}$$

and

$$\|oldsymbol{
abla}^2(\mathscr{S}^{\star}g)\|_{\Omega} \leq 2rac{\omega}{c_{\mathsf{min}}}\left(1+\mathscr{C}_{\mathrm{st}}
ight)\|g\|_{\mu,\Omega} \hspace{3mm} orall g\in L^2(\Omega).$$

Guaranteed bound

$$\mathscr{C}_{\mathrm{ap}} \leq 2\left(1+\mathscr{C}_{\mathcal{T},\mathrm{i}}
ight) rac{\omega h}{c_{\min}} \mathscr{C}_{\mathrm{st}}$$

Takeaways

The estimator η needs to be "pre-factored" by $C_{\rm st}$ or $C_{\rm ap}$. The "qualitative" behaviors of both quantities are relatively well known.

The behaviour of $\mathscr{C}_{\rm st}$ is only dictated by the PDE. Explicit bounds are available for non-trapping star-shaped obstables.

The approximation factor \mathscr{C}_{ap} depends on the PDE and V_h . When $\mathbf{A} = \mathbf{I}$, Ω is convex and \mathscr{C}_{st} is known, we can bound it nicely.

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Propagation of a plane wave

We consider the propagation of a plane wave in $\Omega = (-1,1)^2$

$$\begin{cases} -\omega^2 \boldsymbol{u} - \Delta \boldsymbol{u} &= 0 \quad \text{in } \Omega, \\ \boldsymbol{\nabla} \boldsymbol{u} \cdot \boldsymbol{n} - i\omega \boldsymbol{u} &= \boldsymbol{g} \quad \text{on } \Gamma_{\mathrm{A}}, \end{cases}$$

where

$$g := \nabla \xi_{\theta} \cdot \boldsymbol{n} - i\omega \xi_{\theta}, \quad \xi_{\theta} := e^{i\omega \, \boldsymbol{d} \cdot \boldsymbol{x}},$$

with $d := (\cos \theta, \sin \theta)$ and $\theta = \pi/12$. The solution is $u = \xi_{\theta}$.



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Plane wave experiment p = 1 and $\omega = \pi$



$$egin{aligned} & \mathcal{E}_{ ext{fem}} := \| \| e_h \| \|_{\omega,\Omega} \ & \mathcal{E}_{ ext{est}} := \eta \ & \widetilde{\mathcal{E}}_{ ext{est}} := (1 + \mathscr{C}_{ ext{ap}}) \eta \end{aligned}$$

Plane wave experiment p = 1 and $\omega = 4\pi$



$$egin{aligned} & \mathcal{E}_{ ext{fem}} := \left\| \left| e_h
ight|
ight|_{\omega,\Omega} \ & \mathcal{E}_{ ext{est}} := \eta \ & \widetilde{\mathcal{E}}_{ ext{est}} := (1 + \mathscr{C}_{ ext{ap}})\eta \end{aligned}$$

Plane wave experiment p = 1 and $\omega = 10\pi$



 $\widetilde{E}_{\text{est}} := (1 + \mathscr{C}_{\text{ap}})\eta$

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Plane wave experiment p = 1 and $\omega = 20\pi$



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Plane wave experiment p = 4 and $\omega = 10\pi$



 $egin{aligned} & \mathcal{E}_{ ext{fem}} := \| \| oldsymbol{e}_h \| \|_{\omega,\Omega} \ & \mathcal{E}_{ ext{est}} := \eta \ & \widetilde{\mathcal{E}}_{ ext{est}} := (1 + \mathscr{C}_{ ext{ap}})\eta \end{aligned}$

Plane wave experiment p = 4 and $\omega = 20\pi$





Plane wave experiment p = 4 and $\omega = 40\pi$



$$egin{aligned} & \mathcal{F}_{ ext{fem}} := \|\!|\!| e_h |\!|\!|_{\omega,\Omega} \ & \mathcal{E}_{ ext{est}} := \eta \ & \widetilde{\mathcal{E}}_{ ext{est}} := (1 + \mathscr{C}_{ ext{ap}})\eta \end{aligned}$$

Plane wave experiment p = 4 and $\omega = 60\pi$



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Scattering by an non-trapping obstacle

We now consider a scattering problem

$$\begin{cases} -\omega^2 u - \Delta u &= 0 \quad \text{in } \Omega, \\ u &= 0 \quad \text{on } \Gamma_{\mathrm{D}}, \\ \nabla u \cdot \mathbf{n} - i\omega u &= g \quad \text{on } \Gamma_{\mathrm{A}}, \end{cases}$$

where again $\boldsymbol{g} = \boldsymbol{\nabla} \xi_{\theta} \cdot \boldsymbol{n} - \boldsymbol{i} \omega \xi_{\theta}$.

We fix the wavenumber $\omega = 10\pi$ and employ \mathbb{P}_3 elements.

We consider a sequence of meshes that are adaptively refined using $\eta_{\mathcal{K}}$.





Solution of the scattering problem



Real (left) and imaginary (right) parts of the solution

(ma)

Estimated error in mesh #1



Estimator η_{K} (left) and elementwise error $\| e_{h} \|_{\omega,K}$ (right)

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Estimated error in mesh #2



Estimator η_{K} (left) and elementwise error $\| e_{h} \|_{\omega,K}$ (right)

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Estimated error in mesh #3



Estimator η_{K} (left) and elementwise error $\| e_{h} \|_{\omega_{K}}$ (right)

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Behavior of the estimator through the adaptive procedure





Behaviors of the estimated and analytical errors in the adaptive procedure

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Conclusions and outlook

We construct an a posteriori error estimator η via flux equilibration. It directly provides guaranteed error estimates at low frequencies.

For high frequencies, η has to be pre-factored, either by \mathscr{C}_{st} or by \mathscr{C}_{ap} . The estimates are asymptotically constant-free. In specific situations, we can provide guaranteed bounds on \mathscr{C}_{st} and \mathscr{C}_{ap} .

There is still a long way toward fully reliable error estimation for high-frequency problems!

T. Chaumont-Frelet, A. Ern, and M. Vohralík, *Numer. Math.*, 2021.

Thank you for your attention!



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