# SECOND-ORDER ABSORBING BOUNDARY CONDITIONS FOR THE WAVE EQUATION: A SOLUTION FOR THE CORNER PROBLEM* 

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#### Abstract

The treatment of domains with corners in the problem of absorbing boundary conditions for the wave equation is very important from a practical point of view. A technical difficulty appears as soon as conditions of order greater than or equal to 2 are considered. A solution is proposed for the two-dimensional case when second-order conditions are used. This solution consists of prescribing an adequate corner condition. The problem thus obtained is analyzed theoretically and the condition is proved to be optimal. The results obtained here are illustrated by numerical simulations. Some extensions to higher-space dimensions and higher-order conditions are proposed.


Key words. wave equations, absorbing boundary conditions, domains with a corner
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1. Introduction. In [4] and [5], Engquist and Majda introduced a sequence of absorbing boundary conditions for the two-dimensional wave equation in the half-plane $\left\{\left(x_{1}, x_{2}\right), x_{2}<0\right\}$ :

$$
\begin{equation*}
\frac{\partial^{2} u}{\partial t^{2}}-\left(\frac{\partial^{2} u}{\partial x_{1}^{2}}+\frac{\partial^{2} u}{\partial x_{2}^{2}}\right)=0 \tag{1.1}
\end{equation*}
$$

(the velocity is supposed to be constant and taken equal to 1 ).
These conditions are given recursively by

$$
\begin{align*}
& \mathbf{B}_{1} u=\frac{\partial u}{\partial t}+\frac{\partial u}{\partial x_{2}}=0 \quad \text { on } \Gamma=\left\{\left(x_{1}, 0\right)\right\}, \\
& \mathbf{B}_{2} u=\frac{\partial^{2} u}{\partial t^{2}}+\frac{\partial^{2} u}{\partial t \partial x_{2}}-\frac{1}{2} \frac{\partial^{2} u}{\partial x_{1}^{2}}=0 \quad \text { on } \Gamma,  \tag{1.2}\\
& \mathbf{B}_{n+1} u=\frac{\partial}{\partial t}\left(\mathbf{B}_{n} u\right)-\frac{1}{4} \frac{\partial^{2}}{\partial x_{1}^{2}}\left(\mathbf{B}_{n-1} u\right)=0 \quad \text { on } \Gamma .
\end{align*}
$$

In [5] it is shown that, according to the Kreiss criterion [8], each of these boundary conditions is strongly well posed for the wave equation.

If we assume that the interior equation (1.1) also holds on the boundary we may write (cf. [6])

$$
\begin{equation*}
\mathbf{B}_{n} u=\frac{1}{2^{n-1}}\left(\frac{\partial}{\partial t}+\frac{\partial}{\partial x_{2}}\right)^{n} u, \quad n=1,2, \cdots, \tag{1.3}
\end{equation*}
$$

and then it is easy to see that each $\mathbf{B}_{n}$ is transparent for plane-harmonic waves striking

[^0]the boundary at normal incidence. Furthermore, if $u$ is a plane harmonic wave of unit amplitude striking the boundary at an angle of incidence $\theta$ from the normal, then the amplitude $R_{n}$ of the reflected wave (i.e., the reflection coefficient for $\mathbf{B}_{n}$ ) is given by
$$
R_{n}=\left(\frac{1-\cos \theta}{1+\cos \theta}\right)^{n}=o\left(\theta^{2 n}\right), \quad n=1,2, \cdots
$$

We mention that some extensions of conditions (1.2) and (1.3) have been proposed in [7] and [15].

For many practical problems we are interested in restricting numerical calculations, which should be done in theory on an infinite domain, to a bounded domain. Thus an absorbing condition is needed for the boundary of the bounded domain.

For a domain with a smooth boundary, a solution has been proposed by Engquist and Majda in [5]. This solution, which makes use of the theory of pseudodifferential operators, generalizes conditions (1.2). However, if we use a finite-difference scheme on a uniform grid to compute the numerical solution, it seems more interesting to restrict the effective calculations to a rectangle. For simplicity of exposition, we restrict our attention to the case of the quarter-plane $\Omega=\left\{x=\left(x_{1}, x_{2}\right) / x_{1}<0, x_{2}<0\right\}$ and then introduce two absorbing boundaries:

$$
\Gamma_{1}=\left\{\left(x_{1}, 0\right), x_{1} \leqq 0\right\}, \quad \Gamma_{2}=\left\{\left(0, x_{2}\right), x_{2} \leqq 0\right\} .
$$

We will denote by zero the corner point ( $x_{1}=0, x_{2}=0$ ); see below.


The initial boundary value problem obtained by taking (1.1) in $\Omega$ with the first-order boundary conditions

$$
\begin{array}{ll}
\frac{\partial^{2} u}{\partial t^{2}}-\Delta u=0 & \text { in } \Omega, \\
\frac{\partial u}{\partial t}+\frac{\partial u}{\partial x_{2}}=0 & \text { on } \Gamma_{1},  \tag{1.4}\\
\frac{\partial u}{\partial t}+\frac{\partial u}{\partial x_{1}}=0 & \text { on } \Gamma_{2},
\end{array}
$$

and appropriate initial conditions

$$
\begin{array}{cc}
u(x, 0)=u_{0}(x) & \text { in } \Omega, \\
\frac{\partial u}{\partial t}(x, 0)=u_{1}(x) & \text { in } \Omega \tag{1.5}
\end{array}
$$

is well posed, as we have the following energy identity:

$$
\begin{equation*}
\frac{\partial}{\partial t}\left(\frac{1}{2} \int_{\Omega}\left|\frac{\partial u}{\partial t}\right|^{2} d x+\frac{1}{2} \int_{\Omega}|\nabla u|^{2} d x\right)+\int_{\Gamma}\left|\frac{\partial u}{\partial t}\right|^{2} d \sigma=0 \tag{1.6}
\end{equation*}
$$

( $d \sigma$ denotes the superficial measure on $\Gamma$ ). Thus the presence of the corner does not pose any specific difficulty. However, as indicated by the reflection coefficients and verified by numerical experiments, these first-order conditions give rise to significantly larger reflected waves than do those of second or higher order.

If we look at the second-order conditions written as

$$
\begin{array}{ll}
\frac{\partial^{2} u}{\partial t^{2}}+\frac{\partial^{2} u}{\partial x_{2} \partial t}-\frac{1}{2} \frac{\partial^{2} u}{\partial x_{1}^{2}}=0 & \text { on } \Gamma_{1}, \\
\frac{\partial^{2} u}{\partial t^{2}}+\frac{\partial^{2} u}{\partial x_{1} \partial t}-\frac{1}{2} \frac{\partial^{2} u}{\partial x_{2}^{2}}=0 & \text { on } \Gamma_{2}, \tag{1.7}
\end{array}
$$

it is not at all clear a priori whether the corresponding initial boundary value problem is well posed. Numerically, if we use a classical finite-difference scheme (for example, see [12], [13]), a naive matching of conditions (1.7) at zero is ruled out by the appearance of the second-order spatial derivatives in any expression for $\mathbf{B}_{2}$. Moreover, it can be shown, as we will see in $\S 3$, that there is no unique solution of the problem in the class of finite energy functions if no condition at the corner is specified. Furthermore, it is not clear what corner condition should be chosen, and as reported in [5], an improper choice of the corner condition may generate instabilities.

Thus, what we propose to do here is to introduce a boundary condition at corner zero and to study this condition from both a mathematical and a numerical point of view.

The outline of this paper is as follows. In § 2 we explain how we construct our corner condition and present different extensions of the method we use. In § 3 we analyse theoretically a family of corner conditions, depending on a parameter, containing our condition and also the one proposed previously in [5]. We try to show in what sense our condition is the best one. In § 4 we present various numerical results illustrating the theoretical results of $\S 3$.
2. Derivation of the boundary condition at the corner. The construction of our corner condition is guided by two quite simple principles:
(i) The corner condition should not introduce a singularity; i.e., regular initial data should yield a regular solution.
(ii) The expression of the corner condition should contain only first-order spatial derivatives in a given direction.

We consider the wave equation (1.1) in $\Omega$ with boundary conditions (1.7) and initial conditions (1.5), and we suppose that the initial data functions $u_{0}$ and $u_{1}$ belong to $C_{0}^{\infty}(\Omega)$. The idea is that if $u$ is a $C^{\infty}$ solution then (1.1) and (1.7) should be satisfied at point zero for all time $t$ :

$$
\begin{array}{lll}
\frac{\partial^{2} u}{\partial t^{2}}-\Delta u=0 & \text { at } x_{1}=x_{2}=0 & \text { for } t \geqq 0 \\
\frac{\partial^{2} u}{\partial t^{2}}+\frac{\partial^{2} u}{\partial x_{2} \partial t}-\frac{1}{2} \frac{\partial^{2} u}{\partial x_{1}^{2}}=0 & \text { at } x_{1}=x_{2}=0 & \text { for } t \geqq 0 \\
\frac{\partial^{2} u}{\partial t^{2}}+\frac{\partial^{2} u}{\partial x_{1} \partial t}-\frac{1}{2} \frac{\partial^{2} u}{\partial x_{2}^{2}}=0 & \text { at } x_{1}=x_{2}=0 & \text { for } t \geqq 0
\end{array}
$$

Adding the second two equations and subtracting $\frac{1}{2}$ times the first equation from the resulting sum, we obtain

$$
\frac{3}{2} \frac{\partial^{2} u}{\partial t^{2}}+\frac{\partial^{2} u}{\partial x_{1} \partial t}+\frac{\partial^{2} u}{\partial x_{2} \partial t}=0 \quad \text { at } x_{1}=x_{2}=0 \quad \text { for } t \geqq 0 .
$$

Then, integrating once with respect to time, we arrive at

$$
\begin{equation*}
\frac{3}{2} \frac{\partial u}{\partial t}+\frac{\partial u}{\partial x_{1}}+\frac{\partial u}{\partial x_{2}}=0 \quad \text { at } x_{1}=x_{2}=0 \quad \text { for } t \geqq 0 . \tag{2.1}
\end{equation*}
$$

It is interesting to compare this condition with that proposed by Engquist and Majda in [5]:

$$
\begin{equation*}
\sqrt{2} \frac{\partial u}{\partial t}+\frac{\partial u}{\partial x_{1}}+\frac{\partial u}{\partial x_{2}}=0 \quad \text { at } x_{1}=x_{2}=0 \quad \text { for } t \geqq 0, \tag{2.2}
\end{equation*}
$$

which is transparent for plane-harmonic waves propagating along the diagonal. In practice, conditions (2.1) and (2.2) differ very little as $\sqrt{2} \approx \frac{3}{2}$, although they are obtained from very different considerations.

At this point, as condition (2.1) is necessarily satisfied by any $C^{2}$ solution of (1.1) and (1.7), it is reasonable to ask if this condition serves any purpose other than numerical ones, for which it differs negligibly from (2.2). We will see its theoretical utility in §3.
(i) Generalization to the $n$-dimensional case. Let $n \geqq 2$ and put

$$
\Omega_{n}=\left\{x=\left(x_{1}, x_{2}, \cdots, x_{n}\right) \in \mathbb{R}^{n} ; x_{i}<0 \text { for } j=1,2, \cdots, n\right\} .
$$

We consider the wave equation

$$
\begin{equation*}
\frac{\partial^{2} u}{\partial t^{2}}-\Delta u=0 \quad \text { in } \Omega \quad \text { for } t>0 \tag{2.3}
\end{equation*}
$$

with initial conditions

$$
\begin{array}{cc}
u(x, 0)=u_{0}(x) & \text { for } x \in \Omega, \\
\frac{\partial u}{\partial t}(x, 0)=u_{1}(x) & \text { for } x \in \Omega,
\end{array}
$$

and the second-order absorbing boundary conditions

$$
\begin{equation*}
\frac{\partial^{2} u}{\partial t^{2}}+\frac{\partial^{2} u}{\partial t \partial x_{k}}-\frac{1}{2} \sum_{1 \leqq j \neq k \leqq n} \frac{\partial^{2} u}{\partial x_{j}} \text { on } \Gamma_{k}, \tag{2.4}
\end{equation*}
$$

$k=1, \cdots, n$, where $\Gamma_{k}$ is the boundary face of $\Omega$ whose normal is $x_{k}$. We suppose that $u_{0}$ and $u_{1}$ belong to $C_{0}^{\infty}\left(\Omega_{n}\right)$. If $u$ is a $C^{\infty}$ solution then (2.3) also holds on $\Gamma_{k}$, and the second-order conditions may be written in the form indicated by (1.3):

$$
\begin{equation*}
\left(\frac{\partial}{\partial t}+\frac{\partial}{\partial x_{k}}\right)^{2} u=0 \quad \text { on } \Gamma_{k}, \tag{2.5}
\end{equation*}
$$

$k=1, \cdots, n$.
For $n>2$ in addition to the condition at the corner itself we need a condition for each lower-dimensional face $\Gamma_{J}$ of $\Omega$, where, for $J \subset\{1,2, \cdots, n\}$, we mean by $\Gamma_{J}$

$$
\Gamma_{J}=\left\{x \in \mathbb{R}^{n} ; x_{k}<0 \text { for } k \notin J, x_{k}=0 \text { for } k \in J\right\} .
$$

If $u$ is a $C^{\infty}$ solution then (2.3) as well as (2.5), for all $k \in J$, holds on $\Gamma_{J}$. If we sum these equations and divide by 2 we obtain

$$
\begin{equation*}
\frac{(j+1)}{2} \frac{\partial^{2} u}{\partial t^{2}}+\frac{\partial}{\partial t} \sum_{k \in J} \frac{\partial u}{\partial x_{k}}-\frac{1}{2} \sum_{k \notin J} \frac{\partial u}{\partial x_{k}} \text { on } \Gamma_{J}, \tag{2.6}
\end{equation*}
$$

where $j$ is the number of elements in $J$.
For the corner itself, where $J=\{1,2, \cdots, n\}$, after integration in time, (2.6) becomes

$$
\begin{equation*}
\frac{n+1}{2} \frac{\partial u}{\partial t}+\sum_{j=1}^{n} \frac{\partial u}{\partial x_{j}}=0 \tag{2.7}
\end{equation*}
$$

We note that, by the Engquist-Majda criterion, the corner condition would be

$$
\begin{equation*}
\sqrt{n} \frac{\partial u}{\partial t}+\sum_{j=1}^{n} \frac{\partial u}{\partial x_{j}}=0 \tag{2.8}
\end{equation*}
$$

and that, for sufficiently large $n,(2.7)$ is really quite different from (2.8).
(ii) Generalization to the third-order boundary condition. Again in dimension 2, we consider the third-order boundary conditions (1.2)

$$
\begin{array}{ll}
\frac{\partial^{3} u}{\partial t^{3}}+\frac{\partial^{3} u}{\partial t^{2} \partial x_{2}}-\frac{3}{4} \frac{\partial^{3} u}{\partial t \partial x_{1}^{2}}-\frac{1}{4} \frac{\partial^{3} u}{\partial x_{1}^{2} \partial x_{2}}=0 & \text { on } \Gamma_{1} \\
\frac{\partial^{3} u}{\partial t^{3}}+\frac{\partial^{3} u}{\partial t^{2} \partial x_{1}}-\frac{3}{4} \frac{\partial^{3} u}{\partial t \partial x_{2}^{2}}-\frac{1}{4} \frac{\partial^{3} u}{\partial x_{1}^{2} \partial x_{2}}=0 & \text { on } \Gamma_{2} \tag{2.9}
\end{array}
$$

As before, we assume that $u$ is $C^{\infty}$ so that (1.1) also holds on $\Gamma_{k}, k=1,2$, and we again use the form indicated by (1.3) for the boundary conditions:

$$
\begin{equation*}
\left(\frac{\partial}{\partial t}+\frac{\partial}{\partial x_{k}}\right)^{3} u=0 \quad \text { on } \Gamma_{k}, \tag{2.10}
\end{equation*}
$$

for $k=1,2$.
We sum the following five equations: (2.10) for $k=1$, (2.10) for $k=2,3$ times (1.1) differentiated once with respect to time, (1.1) differentiated once with respect to $x_{1}$, and (1.1) differentiated once with respect to $x_{2}$. Thus we obtain

$$
\begin{equation*}
5 \frac{\partial^{3} u}{\partial t^{3}}+4\left(\frac{\partial^{3} u}{\partial t^{2} \partial x_{1}}+\frac{\partial^{3} u}{\partial t^{2} \partial x_{2}}\right)-\frac{\partial^{3} u}{\partial x_{1}^{2} \partial x_{2}}-\frac{\partial^{3} u}{\partial x_{2}^{2} \partial x_{1}}=0 \quad \text { at } 0 . \tag{2.11}
\end{equation*}
$$

The last two terms on the left-hand side need to be eliminated. We differentiate (2.10) on $\Gamma_{1}$ with respect to $x_{2}$ and (2.10) on $\Gamma_{2}$ with respect to $x_{1}$. Equation (1.1) is differentiated twice, once with respect to each spatial variable. The resulting three equations are summed to obtain an equation that can be integrated in time. We then have

$$
\begin{equation*}
\frac{\partial^{3} u}{\partial t^{2} \partial x_{1}}+\frac{\partial^{3} u}{\partial t^{2} \partial x_{2}}+\frac{\partial^{3} u}{\partial t \partial x_{1} \partial x_{2}}+3\left(\frac{\partial^{3} u}{\partial x_{1}^{2} \partial x_{2}}+\frac{\partial^{3} u}{\partial x_{2}^{2} \partial x_{1}}\right)=0 . \tag{2.12}
\end{equation*}
$$

Summing (2.12) with 3 times (2.11) and integrating the result once in time, we arrive at the corner condition

$$
\begin{equation*}
15 \frac{\partial^{2} u}{\partial t^{2}}+13\left(\frac{\partial^{2} u}{\partial t \partial x_{1}}+\frac{\partial^{2} u}{\partial t \partial x_{2}}\right)+7 \frac{\partial^{2} u}{\partial x_{1} \partial x_{2}}=0 \quad \text { at } 0 . \tag{2.13}
\end{equation*}
$$

As in the case of the second-order boundary condition, the corner condition (2.13) can be extended with no specific difficulty to the wave equation in $\mathbb{R}^{n}$.

It would also be natural to generalize this corner condition to higher-order absorbing boundary conditions (1.2). This problem does not appear to be so easy. For example, as fourth-order spatial derivatives in a given direction appear in the fourthand fifth-order absorbing boundary conditions (see [5]), our conjecture is that three corner conditions are needed, and more generally that $n$ corner conditions are needed for the $2 n$ th- and $(2 n+1)$ st-order conditions.
3. Mathematical analysis of the two-dimensional problem. Our goal in this section is to study from a mathematical point of view the following problem:

$$
\begin{array}{lll}
\frac{\partial^{2} u}{\partial t^{2}}-\Delta u=0 & \text { in } \Omega & \text { for } t>0 \\
\frac{\partial^{2} u}{\partial t^{2}}+\frac{\partial^{2} u}{\partial x_{2} \partial t}-\frac{1}{2} \frac{\partial^{2} u}{\partial x_{1}^{2}}=0 & \text { on } \Gamma_{1} & \text { for } t>0 \\
\frac{\partial^{2} u}{\partial t^{2}}+\frac{\partial^{2} u}{\partial x_{1} \partial t}-\frac{1}{2} \frac{\partial^{2} u}{\partial x_{2}^{2}}=0 & \text { on } \Gamma_{2} & \text { for } t>0
\end{array}
$$

( $\mathrm{P}_{\gamma}$ )

$$
\begin{array}{ll}
\gamma \frac{\partial u}{\partial t}+\frac{\partial u}{\partial x_{1}}+\frac{\partial u}{\partial x_{2}}=0 & \text { at } 0 \quad \text { for } t>0 \\
u(x, 0)=u_{0}(x) & \text { in } \Omega, \\
\frac{\partial u}{\partial t}(x, 0)=u_{1}(x) & \text { in } \Omega .
\end{array}
$$

For this problem we have the second-order absorbing boundary conditions:

$$
\begin{array}{lll}
\frac{\partial^{2} u}{\partial t^{2}}+\frac{\partial^{2} u}{\partial x_{2} \partial t}-\frac{1}{2} \frac{\partial^{2} u}{\partial x_{1}^{2}}=0 & \text { on } \Gamma_{1} & \text { for } t>0 \\
\frac{\partial^{2} u}{\partial t^{2}}+\frac{\partial^{2} u}{\partial x_{1} \partial t}-\frac{1}{2} \frac{\partial^{2} u}{\partial x_{2}^{2}}=0 & \text { on } \Gamma_{2} & \text { for } t>0 \tag{CL2}
\end{array}
$$

and we consider a general corner condition:

$$
\gamma \frac{\partial u}{\partial t}+\frac{\partial u}{\partial x_{1}}+\frac{\partial u}{\partial x_{2}}=0 \quad \text { at } 0 \quad \text { for } t>0
$$

where $\gamma$ denotes a positive parameter. This generalizes condition (2.1) derived in § 2, and condition (2.2) of Engquist and Majda. We put

$$
\gamma^{*}=\frac{3}{2}
$$

It is not obvious how to obtain an existence and uniqueness result for $\left(\mathrm{P}_{\gamma}\right)$ via classical methods based on energy identities and a priori estimates. Moreover, this problem does not fall into the category of corner problems previously treated in the literature (cf. [9], [11], and [14]). Thus we develop an analysis specific to this problem. Nonetheless, we cite the work of Lemrabet [10] on the analogous problem for the Ventcel boundary conditions.

It is clear that $\left(\mathrm{P}_{\gamma}\right)$ can have a $C^{\infty}\left(\bar{\Omega} \times \mathbb{R}^{+}\right)$solution only if $\gamma=\gamma^{*}$, and that in this case condition ( $\mathrm{CC} \gamma$ ) is superfluous. In $\S 3.3$ we will see that if the initial data functions $u_{0}$ and $u_{1}$ belong to $C_{0}^{\infty}(\Omega)$, then there does indeed exist a solution $u$ of ( $\mathrm{P}_{\gamma^{*}}$ ) in $C^{\infty}\left(\bar{\Omega} \times \mathbb{R}^{+}\right)$. (The corner condition is not used for this construction.)

In $\S 3.1$ the corner condition is used to derive a weak formulation of $\left(\mathrm{P}_{\gamma}\right)$. This weak formulation is used in $\S 3.2$ to obtain a uniqueness result and in $\S 3.3$ to extend the existence result for $\gamma=\gamma^{*}$ to the case in which the initial data lie in $H^{5} \times H^{4}$.

In § 3.4 we obtain an existence result for arbitrary $\gamma>0$ and point out the existence of a singular corner wave when $\gamma \neq \gamma^{*}$. We then analyze the properties of this corner wave.
3.1. Weak formulation of the problem. Suppose there exists a function $u$ such that

$$
\begin{aligned}
& u \in \mathbf{C}^{3}\left(\bar{\Omega}^{*} \times \mathbb{R}^{+}\right) \cap \mathbf{C}^{1}(\bar{\Omega} \times[0, \infty)), \\
& \forall t \geqq 0 \quad u(t) \text { has compact support, } \\
& u \text { satisfies the equations of }\left(\mathrm{P}_{\gamma}\right) .
\end{aligned}
$$

We differentiate the wave equation (1.1) with respect to $t$, multiply by a test function $v\left(x_{1}, x_{2}\right)$ in $C_{0}^{\infty}(\bar{\Omega})$, and integrate in space to see that such a function $u$ satisfies

$$
\frac{d^{3}}{d t^{3}}\left(\int_{\Omega} u v d x\right)-\int_{\Omega} \Delta \frac{\partial u}{\partial t} v d x=0
$$

Using Green's formula we obtain

$$
\begin{equation*}
\frac{d^{3}}{d t^{3}}\left(\int_{\Omega} u(t) v d x\right)+\frac{d}{d t}\left(\int_{\Omega} \nabla u(t) \nabla v d x\right)-\int_{\Gamma} \frac{\partial^{2} u}{\partial n \partial t}(t) v d \sigma=0 . \tag{3.2}
\end{equation*}
$$

Let us transform the boundary integral as follows:

$$
-\int_{\Gamma} \frac{\partial^{2} u}{\partial n \partial t}(t) v d \sigma=-\int_{\Gamma_{1}} \frac{\partial^{2} u}{\partial x_{2} \partial t} v d x_{1}-\int_{\Gamma_{2}} \frac{\partial^{2} u}{\partial x_{1} \partial t} v d x_{2} .
$$

Using the boundary conditions on $\Gamma_{1}$ and $\Gamma_{2}$ we see that

$$
\begin{align*}
-\int_{\Gamma} \frac{\partial^{2} u}{\partial n \partial t} v d \sigma=\frac{d^{2}}{d t^{2}} & \left(\int_{\Gamma_{1}} u v d x_{1}\right)-\frac{1}{2} \int_{\Gamma_{2}} \frac{\partial^{2} u}{\partial x_{1}^{2}} v d x_{1} \\
& +\frac{d^{2}}{d t^{2}}\left(\int_{\Gamma_{2}} u v d x_{2}\right)-\frac{1}{2} \int_{\Gamma_{2}} \frac{\partial^{2} u}{\partial x_{2}^{2}} v d x_{2} \tag{3.3}
\end{align*}
$$

A double integration by parts leads to the equality

$$
\begin{gather*}
-\frac{1}{2} \int_{\Gamma_{1}} \frac{\partial^{2} u}{\partial x_{1}^{2}} v d x_{1}-\frac{1}{2} \int_{\Gamma_{2}} \frac{\partial^{2} u}{\partial x_{2}^{2}} v d x_{2}=\frac{1}{2} \int_{\Gamma_{1}} \frac{\partial u}{\partial x_{1}} \frac{\partial v}{\partial x_{1}} d x_{1}+\frac{1}{2} \int_{\Gamma_{2}} \frac{\partial u}{\partial x_{2}} \frac{\partial v}{\partial x_{2}} d x_{2}  \tag{3.4}\\
-\frac{1}{2}\left\{\frac{\partial u}{\partial x_{1}}(0)+\frac{\partial u}{\partial x_{2}}(0)\right\} v(0)
\end{gather*}
$$

We finally use the corner equation ( $\mathrm{CC} \gamma$ ) to obtain

$$
\begin{align*}
-\frac{1}{2} \int_{\Gamma_{1}} \frac{\partial^{2} u}{\partial x_{1}^{2}} v d x_{1}-\frac{1}{2} \int_{\Gamma_{2}} \frac{\partial^{2} u}{\partial x_{2}^{2}} v d x_{2}=\frac{1}{2} & \left\{\int_{\Gamma_{1}} \frac{\partial u}{\partial x_{1}} \frac{\partial v}{\partial x_{1}} d x_{1}+\int_{\Gamma_{2}} \frac{\partial u}{\partial x_{2}} \frac{\partial v}{\partial x_{2}} d x_{2}\right\}  \tag{3.5}\\
+ & \frac{\gamma}{2} \frac{d}{d t}(u(0) v(0)) .
\end{align*}
$$

Thus, regrouping (3.2), (3.3), and (3.5), we have proved that $u$ satisfies the equation

$$
\begin{align*}
\frac{d^{3}}{d t^{3}}\left(\int_{\Omega} u v d x\right)+\frac{d^{2}}{d t^{2}} & \left(\int_{\Gamma} u v d \sigma\right)+\frac{d}{d t}\left(\int_{\Omega} \nabla u \nabla v d x+\frac{\gamma}{2} u(0) v(0)\right)  \tag{3.6}\\
& +\frac{1}{2} \int_{\Gamma} \frac{\partial u}{\partial \tau} \frac{\partial v}{\partial \tau} d \sigma=0 \quad \text { for all } v \in \mathbf{C}_{0}^{\infty}(\bar{\Omega}) .
\end{align*}
$$

It is clear that (3.6) can have a sense even if $u$ is less regular than required in (3.1). Thus we are led to introduce the functional space

$$
\begin{align*}
& V=\left\{v \in H^{1}(\Omega) ; v_{1}\left(x_{1}\right)=v\left(x_{1}, 0\right) \in H^{1}\left(\Gamma_{1}\right), v_{2}\left(x_{2}\right)=v\left(0, x_{2}\right) \in H^{1}\left(\Gamma_{2}\right),\right. \\
& v_{1}(0)=v_{2}(0)\stackrel{\text { def }}{=} v(0)\} . \tag{3.7}
\end{align*}
$$

Note that $V$ is well defined. Indeed, for $v$ in $H^{1}(\Omega)$ the trace $v_{1}$ (respectively, $v_{2}$ ) belongs to $H^{1 / 2}\left(\Gamma_{1}\right)$ (respectively, $H^{1 / 2}\left(\Gamma_{2}\right)$ ), and thus we can ask if it also belongs to $H^{1}\left(\Gamma_{1}\right)$ (respectively, $\left.H^{1}\left(\Gamma_{2}\right)\right)$. Moreover, $v_{j}(0)$ makes sense as $v_{j}$ belongs to $H^{1}((-\infty, 0))$. Note also that $V$ contains $\mathbf{C}_{0}^{\infty}(\bar{\Omega})$. We define the following norm on $V$ :

$$
\begin{equation*}
\|v\|_{V}^{2}=\int_{\Omega}|v|^{2} d x+\int_{\Omega}|\nabla v|^{2} d x+\int_{\Gamma}\left|\frac{\partial v}{\partial \tau}\right|^{2} d \sigma+|v(0)|^{2} \tag{3.8}
\end{equation*}
$$

( $\partial v / \partial \tau$ is the function on $L^{2}(\Gamma)$ defined by $\left.\partial v / \partial \tau \mid \Gamma_{j}=\partial v_{j} / \partial x_{j}, j=1,2\right)$. It is easy to show Lemma 3.1.

Lemma 3.1. Equipped with the norm (3.8), V is a Hilbert space.
Remark. Note that $V$ is nothing but the closure of $\mathbf{C}_{0}^{\infty}(\bar{\Omega})$ for the norm $\|\cdot\|_{V}$.
To complete our mathematical framework, we introduce, for $H$ a Hilbert space (with norm $\|\cdot\|_{H}$ ) and $\sigma$ a positive number, the space

$$
\begin{equation*}
L_{\sigma}^{1}\left(\mathbb{R}^{+} ; H\right)=\left\{v(t): \mathbb{R}^{+} \rightarrow H ; \int_{0}^{+\infty}\|v(t)\|_{H} e^{-\sigma t} d t<+\infty\right\} . \tag{3.9}
\end{equation*}
$$

Such a space is of interest because it is possible to define the Laplace transform of any $v$ in $V$ by (see [2], [3])

$$
\begin{equation*}
\hat{v}(p)=\int_{0}^{+\infty} v(t) e^{-p t} d t \quad \text { for } \operatorname{Re}(p)>\sigma \tag{3.10}
\end{equation*}
$$

and to show that the function $p \rightarrow \hat{v}(p) \in H$ is analytic in the half-plane $\operatorname{Re}(p)>\sigma$. In particular we have the property

$$
\begin{array}{ll}
\hat{v}(p)=0 & \text { for all } p \in \mathbb{R} \\
& \text { such that } p>\sigma \text { implies } v(t)=0 \text { for all } t \geqq 0 . \tag{3.11}
\end{array}
$$

Note also that $L_{\sigma}^{1}\left(\mathbb{R}^{+} ; V\right) \subset L_{\text {loc }}^{1}\left(\mathbb{R}^{+} ; V\right) \subset \mathbf{D}^{\prime}\left(\mathbb{R}^{+} ; V\right)$.
Now we can give a definition.
Definition 3.1. A function $u(t)$ is a weak solution of problem $\left(\mathrm{P}_{\gamma}\right)$ if and only if there exists $\sigma>0$ such that

$$
\begin{align*}
& \left(u, \frac{d u}{d t}\right) \in L_{\sigma}^{1}\left(\mathbb{R}^{+} ; V\right) \times L_{\sigma}^{1}\left(\mathbb{R}^{+} ; L^{2}(\Omega)\right),  \tag{i}\\
& u(0)=u_{0} \in V, \quad \frac{d u}{d t}(0)=u_{1} \in L^{2}(\Omega),  \tag{ii}\\
& \frac{d^{3}}{d t^{3}}\left(\int_{\Omega} u v d x\right)+\frac{d^{2}}{d t^{2}}\left(\int_{\Gamma} u v d \sigma\right)+\frac{d}{d t}\left(\int_{\Omega} \nabla u \nabla v d x+\frac{\gamma}{2} u(0) v(0)\right) \\
& \quad+\frac{1}{2} \int_{\Gamma} \frac{\partial u}{\partial \tau} \frac{\partial v}{\partial \tau} d \sigma=0 \text { in } \mathbf{D}^{\prime}\left(\mathbb{R}^{+}\right) \quad \text { for all } v \in V .
\end{align*}
$$

Remarks. In (i), the derivative $d u / d t$ is taken in the sense of distributions with values in $L^{2}(\Omega)$.

If we choose $v \in \mathbf{D}(\Omega)$, it is easy to see that (iii) can be interpreted as

$$
\frac{\partial^{2} u}{\partial t^{2}}-\Delta u=0 \quad \text { in } \mathbf{D}^{\prime}\left(\mathbb{R}^{+} \times \Omega\right)
$$

which shows, as $\Delta \in \mathscr{L}\left(H^{1}(\Omega) ; H^{-1}(\Omega)\right)$, that $d^{2} u / d t^{2}$ belongs to the space $L_{\sigma}^{1}\left(\mathbb{R}^{+}\right.$; $\left.H^{-1}(\Omega)\right)$. Then we deduce that

$$
u \in \mathbf{C}^{1}\left([0,+\infty) ; H^{-1}(\Omega)\right) \cap \mathbf{C}^{0}\left([0,+\infty) ; L^{2}(\Omega)\right)
$$

which permits us to give a meaning to initial conditions (ii).
3.2. A uniqueness result. Let $u$ be a solution of $\left(\mathrm{P}_{\gamma}\right)$ in the sense of Definition 3.1. Then for $\operatorname{Re}(p)>\sigma$ its Laplace transform $\hat{u}(p)$ is, at least formally, a solution of

$$
\begin{array}{ll}
-\Delta \hat{u}+p^{2} \hat{u}=u_{1}+p u_{0}=f & \text { in } \Omega, \\
-\frac{1}{2} \frac{\partial^{2} \hat{u}}{\partial x_{1}^{2}}+p \frac{\partial \hat{u}}{\partial x_{2}}+p^{2} \hat{u}=0 & \text { on } \Gamma_{1}, \\
-\frac{1}{2} \frac{\partial^{2} \hat{u}}{\partial x_{2}^{2}}+p \frac{\partial \hat{u}}{\partial x_{1}}+p^{2} \hat{u}=0 & \text { on } \Gamma_{2}, \\
\frac{\partial \hat{u}}{\partial x_{1}}+\frac{\partial \hat{u}}{\partial x_{2}}+\gamma p \hat{u}=0 & \text { at } 0,
\end{array}
$$

or more exactly, using the weak formulation, of

$$
\begin{align*}
& \hat{u}(p) \in V, \\
& \hat{a}(p, \hat{u}(p), \hat{v})=\langle L(p), \hat{v}\rangle \quad \text { for all } \hat{v} \in V, \tag{3.13}
\end{align*}
$$

where the bilinear form $\hat{a}(p, \cdot, \cdot)$ is defined by

$$
\begin{align*}
\hat{a}(p, \hat{u}, \hat{v})=\int_{\Omega} \nabla & \hat{u} \cdot \overline{\nabla \hat{v}} d x+p^{2} \int_{\Omega} \hat{u} \overline{\hat{v}} d x \\
& \quad+p \int_{\Gamma} \hat{u} \overline{\hat{v}} d \sigma+\frac{1}{2 p} \int_{\Gamma} \frac{\partial \hat{u}}{\partial \tau} \frac{\bar{v}}{\partial \tau} d \sigma+\frac{\gamma}{2} \hat{u}(0) \overline{\hat{v}(0)} \tag{3.14}
\end{align*}
$$

and the linear form $\langle\hat{L}(p), \cdot\rangle$ is given by

$$
\begin{equation*}
\langle\hat{L}(p), \hat{v}\rangle=\int_{\Omega} u_{1} \overline{\hat{v}} d x+p \int_{\Omega} u_{0} \overline{\hat{v}} d x . \tag{3.15}
\end{equation*}
$$

We have the following lemma.
Lemma 3.2. For $\gamma>0$, if $p$ is a positive real number, the bilinear form $\hat{a}(p, \hat{u}, \hat{v})$ is $V$ elliptic.

Proof. It suffices to write

$$
\hat{a}(p, \hat{u}, \hat{u})=\int_{\Omega}|\nabla \hat{u}|^{2} d x+p^{2} \int_{\Omega}|\hat{u}|^{2} d x+p \int_{\Gamma}|\hat{u}|^{2} d \sigma+\frac{1}{2 p} \int_{\Gamma}\left|\frac{\partial \hat{u}}{\partial \tau}\right|^{2} d \sigma+\frac{\gamma}{2}|\hat{u}(0)|^{2}
$$

to deduce that $\hat{a}(p, \hat{u}, \hat{u}) \geqq \inf \left(1,1 / 2 p, p^{2}, \gamma / 2\right)\|\hat{u}\|_{V}^{2}$.
Then, using the Lax-Milgram lemma, we see that problem (3.13) admits, for $\left(u_{0}, u_{1}\right)$ in $V \times L^{2}(\Omega)$, a unique solution $\hat{u}(p)$ in $H$. In particular, if $u_{0}=u_{1}=0$, we necessarily have $\hat{u}(p)=0$ for all $p>\sigma$. Using property (3.9), we deduce the following theorem.

Theorem 3.1. The solution $u$ of $P_{\gamma}$, if it exists, is unique.
Of course, we could try to develop an existence theory for $\left(\mathbf{P}_{\gamma}\right)$ using the Laplace transform method. We could hope to prove an existence and uniqueness result for problem (3.12) when $p=\eta+i \omega, \eta>0$ being fixed and $\omega$ varying in $\mathbb{R}$, and to obtain estimates on $\hat{u}(\eta+i \omega$ ) (in $V$ ) for any $\omega$ in $\mathbb{R}$. We did not succeed in obtaining such estimates (we can show that such estimates are available in the region $\omega^{2} \leqq 3 \eta^{2}$, but we do not know how to extend these estimates to the whole line $p=\eta+i \omega$ ).
3.3. The existence of a smooth solution for $\boldsymbol{\gamma}=\boldsymbol{\gamma}^{*}=\frac{3}{2}$. We assume that the initial data functions $u_{0}$ and $u_{1}$ belong to $C_{0}^{\infty}(\Omega)$. We will construct a solution $u$ of ( $\mathrm{P}_{\gamma^{*}}$ ) of class $C^{\infty}$. Toward this end we remark that if $u$ is such a function then the $C^{\infty}$ function $v$ defined by

$$
\begin{align*}
& v=L_{1} L_{2} u, \\
& L_{1}=\frac{\partial^{2}}{\partial t^{2}}+\frac{\partial^{2}}{\partial x_{2} \partial t}-\frac{1}{2} \frac{\partial^{2}}{\partial x_{1}^{2}},  \tag{3.16}\\
& L_{2}=\frac{\partial^{2}}{\partial t^{2}}+\frac{\partial^{2}}{\partial x_{1} \partial t}-\frac{1}{2} \frac{\partial^{2}}{\partial x_{2}^{2}},
\end{align*}
$$

satisfies

$$
\begin{array}{ll}
\frac{\partial^{2} v}{\partial t^{2}}-\Delta v=0 & \text { in } \Omega \text { for } t>0, \\
v_{\mid \Gamma}=0 & \text { for } t>0,  \tag{3.17}\\
v(x, 0)=v_{0}(x) & \text { in } \Omega, \\
\frac{\partial v}{\partial t}(x, 0)=v_{1}(x) & \text { in } \Omega,
\end{array}
$$

where

$$
\begin{align*}
& v_{0}=\frac{1}{2} \Delta^{2} u_{0}+\frac{1}{4} \frac{\partial^{4} u_{0}}{\partial x_{1}^{2} \partial x_{2}^{2}}+\frac{\partial^{2} \Delta u_{0}}{\partial x_{1} \partial x_{2}}+\frac{1}{2}\left(\frac{\partial^{3} u_{1}}{\partial x_{1}^{3}}+\frac{\partial^{3} u_{1}}{\partial x_{2}^{3}}\right)+\frac{\partial^{3} u_{1}}{\partial x_{1}^{2} \partial x_{2}}+\frac{\partial^{3} u_{1}}{\partial x_{1} \partial x_{2}^{2}}, \\
& v_{1}=\frac{1}{2} \Delta^{2} u_{1}+\frac{1}{4} \frac{\partial^{4} u_{1}}{\partial x_{1}^{2} \partial x_{2}^{2}}+\frac{\partial^{2} \Delta u_{1}}{\partial x_{1} \partial x_{2}}+\left(\frac{\partial}{\partial x_{1}}+\frac{\partial}{\partial x_{2}}\right) \Delta^{2} u_{0}-\frac{1}{2}\left(\frac{\partial^{3}}{\partial x_{1}^{3}}+\frac{\partial^{3}}{\partial x_{2}^{3}}\right) \Delta u_{0} . \tag{3.18}
\end{align*}
$$

Furthermore, as $u$ is a solution of the wave equation, we deduce that

$$
\begin{equation*}
v=L u=\frac{1}{4}\left(\frac{\partial}{\partial t}+\frac{\partial}{\partial x_{1}}\right)^{2}\left(\frac{\partial}{\partial t}+\frac{\partial}{\partial x_{2}}\right)^{2} u . \tag{3.19}
\end{equation*}
$$

Now, to actually construct a $C^{\infty}$ solution of $\left(\mathrm{P}_{\gamma^{*}}\right)$, we let $v$ denote the unique solution of (3.17) whose existence is guaranteed by the theory of images. We then let $u$ be the unique solution in the quarter-plane $\Omega$ of

$$
\begin{align*}
& L u=v \quad \text { in } \Omega \text { for } t>0, \\
& u=u_{0}, \quad \frac{\partial u}{\partial t}=u_{1}, \quad \frac{\partial^{2} u}{\partial t^{2}}=\Delta u_{0}, \quad \frac{\partial^{3} u}{\partial t^{3}}=\Delta u_{1} \quad \text { in } \Omega \quad \text { for } t=0 . \tag{3.20}
\end{align*}
$$

This problem is well posed, as its solution is obtained by integrating in the quarter-plane $\Omega=\left\{\left(x_{1}, x_{2}\right) ; x_{1}<0, x_{2}<0\right\}$. Moreover, its solution $u$ is $C^{\infty}$.

Theorem 3.2. The function $u$ defined by (3.17), (3.18), and (3.20) is a classical solution of $\left(\mathrm{P}_{\gamma^{*}}\right)$.

Proof. To see that $u$ satisfies the wave equation in $\Omega$ it suffices to note that $w=\partial^{2} u / \partial t^{2}-\Delta u$ is the unique solution to

$$
\begin{aligned}
& L w=0 \quad \text { in } \Omega \quad \text { for } t>0, \\
& w=\frac{\partial w}{\partial t}=\frac{\partial^{2} w}{\partial t^{2}}=\frac{\partial^{3} w}{\partial t^{3}}=\frac{\partial^{4} w}{\partial t^{4}}=0 \quad \text { in } \Omega \quad \text { for } t=0 .
\end{aligned}
$$

Thus $w=0$ in $\Omega$.
Since $u$ is $C^{\infty}, u$ satisfies the wave equation on $\Gamma$. Thus to see that (CL1) holds for $u$ on $\Gamma_{1}$ it is sufficient to show that $w_{2}=\left(\partial / \partial t+\partial / \partial x_{2}\right)^{2} u$ vanishes on $\Gamma_{1}$. For this result we note that $g_{1}\left(x_{1}, t\right)=w_{2}\left(x_{1}, 0, t\right)$ is the unique solution of

$$
\begin{array}{ll}
\left(\frac{\partial}{\partial t}+\frac{\partial}{\partial x_{1}}\right)^{2} g_{1}\left(x_{1}, t\right)=0 & \text { for } x_{1} \in \Gamma_{1} \text { and } t>0, \\
g_{1}\left(x_{1}, 0\right)=\frac{\partial}{\partial t} g_{1}\left(x_{1}, 0\right)=0 & \text { for } x_{1} \in \Gamma_{1},
\end{array}
$$

and thus that $g_{1}\left(x_{1}, t\right)=0$ for $x_{1} \in \Gamma_{1}$.
That (CL2) holds for $u$ on $\Gamma_{2}$ is demonstrated in the same manner.
For completeness we would like to check that the solution $u$ constructed above belongs to the class of functions for which we obtained the uniqueness result in § 3.2, i.e., we would like to check that $u$ is a weak solution of $\left(\mathrm{P}_{\gamma}\right)$ in the sense of Definition 3.1. (While we know that $u$ is $C^{\infty}$, we do not yet know that it belongs to the space $L_{\sigma}^{1}\left(\mathbb{R}^{+} ; V\right)$, for which we must show that the energy does not increase more than exponentially in time.)

In checking the solution we obtain estimates that allow us to extend our existence and uniqueness result to the case in which the initial data $\left(u_{0}, u_{1}\right)$ lies in $H^{5}(\Omega) \times H^{4}(\Omega)$. (This result is probably not optimal, though we have not succeeded in improving it.)

Theorem 3.3. The classical solution $u$ of ( $\mathrm{P}_{\gamma^{*}}$ ) defined by (3.17), (3.18), and (3.20) is also a weak solution in the sense of Definition 3.1. Thus it is unique.

Proof. Suppose that $f$ and $g$ denote functions regular in $\Omega$ such that

$$
\frac{\partial f}{\partial t}+\frac{\partial f}{\partial x_{1}}=g \quad \text { in } \Omega .
$$

If we multiply this equation by $f$, integrate over $\Omega$, and then integrate from zero to $t$ we get

$$
\begin{equation*}
\frac{1}{2} \int_{\Omega}|f|^{2} d x+\int_{0}^{t} \int_{\Gamma_{2}}|f|^{2} d x_{2} d s=\frac{1}{2} \int_{\Omega}\left|f_{0}\right|^{2}+\int_{0}^{t} \int_{\Omega} f g d x d s \tag{3.21}
\end{equation*}
$$

where $f_{0}(x) \equiv f(x, 0)$. Then, using Gronwall's Lemma, we obtain

$$
\begin{equation*}
\int_{\Omega}|f|^{2} d x \leqq\left\{\int_{\Omega}\left|f_{0}\right|^{2} d x+\int_{0}^{t} \int_{\Omega}|g|^{2} d x d s\right\} \tag{3.22}
\end{equation*}
$$

and plugging (3.22) into (3.21) we have

$$
\begin{equation*}
\int_{0}^{t} \int_{\Gamma_{2}}|f|^{2} d x_{2} d s \leqq(1+2 t) \int_{\Omega}\left|f_{0}\right|^{2} d x+\int_{0}^{t}(1+t-s)\left(\int_{\Omega}|g|^{2} d x\right) d s \tag{3.23}
\end{equation*}
$$

Now consider two functions $w$ and $g$ defined in $\Omega$ and related by the sequence of equalities (with $w=w_{1122}$ )

$$
\begin{aligned}
& \frac{\partial w_{1}}{\partial t}+\frac{\partial w_{1}}{\partial x_{1}}=g, \\
& \frac{\partial w_{12}}{\partial t}+\frac{\partial w_{12}}{\partial x_{2}}=w_{1}, \\
& \frac{\partial w_{112}}{\partial t}+\frac{\partial w_{112}}{\partial x_{1}}=w_{12}, \\
& \frac{\partial w_{1122}}{\partial t}+\frac{\partial w_{1122}}{\partial x_{2}}=w_{112} .
\end{aligned}
$$

Then assuming that $\int_{\Omega}|g|^{2} d x \leqq\left\|G_{0}\right\|^{2}$ uniformly in time, we get, using (3.22),

$$
\begin{aligned}
& \int\left|w_{1}\right|^{2} d x \leqq 2\left(\left|w_{1}(0)\right|^{2}+\left\|G_{0}\right\|^{2} t\right), \\
& \int\left|w_{12}\right|^{2} d x \leqq 2\left\{\left|w_{12}(0)\right|^{2}+2\left|w_{1}(0)\right|^{2} t+\left\|G_{0}\right\|^{2} \frac{t^{2}}{2}\right\}, \\
& \int\left|w_{112}\right|^{2} d x \leqq 2\left\{\left|w_{112}(0)\right|^{2}+2\left|w_{12}(0)\right|^{2} t+2\left|w_{1}(0)\right|^{2} t^{2}+\|+\| G_{0} \|^{2} \frac{t^{3}}{3}\right\},
\end{aligned}
$$

and finally, with the aid of (3.22) and (3.23), we see that

$$
\begin{align*}
& \int|w|^{2} d x \leqq C\left(1+t^{4}\right)\left(|w(0)|^{2}+\left|w_{112}(0)\right|^{2}+\left|w_{12}(0)\right|^{2}+\left|w_{1}(0)\right|^{2}+\left\|G_{0}\right\|^{2}\right), \\
& \int_{0}^{t} \int_{\Gamma_{2}}|w|^{2} d x_{2} d s \leqq C\left(1+t^{5}\right)\left(|w(0)|^{2}+\left|w_{112}(0)\right|^{2}+\left|w_{12}(0)\right|^{2}+\left|w_{1}(0)\right|^{2}+\left\|G_{0}\right\|^{2}\right) . \tag{3.24}
\end{align*}
$$

We can now invert the roles of $x_{1}$ and $x_{2}$ in the previous estimates. We then introduce $w_{2}, w_{21}$, and $w_{221}$ to obtain

$$
\begin{align*}
& \int|w|^{2} d x \leqq C\left(1+t^{4}\right)\left(|w(0)|^{2}+\left|w_{221}(0)\right|^{2}+\left|w_{21}(0)\right|^{2}+\left|w_{2}(0)\right|^{2}+\left\|G_{0}\right\|^{2}\right) \\
& \int_{0}^{t} \int_{\Gamma_{1}}|w|^{2} d x_{1} \leqq C\left(1+t^{5}\right)\left(|w(0)|^{2}+\left|w_{221}(0)\right|^{2}+\left|w_{21}(0)\right|^{2}+\left|w_{2}(0)\right|^{2}+\left\|G_{0}\right\|^{2}\right) . \tag{3.25}
\end{align*}
$$

Now we can apply estimates (3.24) and (3.25) successively to

$$
\begin{array}{ll}
w=\frac{\partial u}{\partial t}, & g=\frac{\partial v}{\partial t} \\
w=\frac{\partial u}{\partial x_{1}}, & g=\frac{\partial v}{\partial x_{1}} \\
w=\frac{\partial u}{\partial x_{2}}, & g=\frac{\partial v}{\partial v x_{2}}
\end{array}
$$

and we can take $\left\|G_{0}\right\|^{2}=\int\left|v_{1}\right|^{2} d x+\int\left|\nabla v_{0}\right|^{2} d x$ since $v$ satisfies a classical energy identity.

Using the relations between $\left(v_{0}, v_{1}\right)$ and $\left(u_{0}, u_{1}\right)$ it is easy to get the following estimate:

$$
\begin{gathered}
\int_{\Omega}\left|\frac{\partial u}{\partial t}\right|^{2} d x+\int_{\Omega}|\nabla u|^{2} d x+\int_{\Gamma}\left|\frac{\partial u}{\partial t}\right|^{2} d \sigma+\int_{\Gamma}\left|\frac{\partial u}{\partial \tau}\right|^{2} d \sigma+\int_{\Gamma}\left|\frac{\partial u}{\partial \nu}\right|^{2} d \sigma \\
\leqq C\left(1+t^{5}\right)\left(\left\|u_{0}\right\|_{H^{5}(\Omega)}^{2}+\left\|u_{1}\right\|_{H^{4}(\Omega)}^{2}\right),
\end{gathered}
$$

where $\tau$ and $\nu$ denote unit tangential and normal vector fields, respectively, on $\Gamma$. The theorem follows.

Decomposition of the solution. It is interesting to analyze the structure of the solution we constructed with the help of the theory of images. Indeed, the solution $v$ of (3.20) can be written as

$$
v=v_{I}+v_{R}^{1}+v_{R}^{2}+v_{R}^{0},
$$

where ( $v_{I}, v_{R}^{1}, v_{R}^{2}, v_{R}^{0}$ ) are, respectively, the restrictions to $\Omega$ of the solutions of the wave equation in the whole plane $\mathbb{R}^{2}$ corresponding to the following initial data ( $\tilde{v}_{0}, \tilde{v}_{1}$ being the extensions of $v_{0}, v_{1}$ by 0 outside $\Omega$ ):

$$
\begin{array}{ll}
v_{I}\left(x_{1}, x_{2}, 0\right)=\tilde{v}_{0}\left(x_{1}, x_{2}\right), & \frac{\partial v_{I}}{\partial t}\left(x_{1}, x_{2}, 0\right)=\tilde{v}_{1}\left(x_{1}, x_{2}\right), \\
v_{R}^{1}\left(x_{1}, x_{2}, 0\right)=-\tilde{v}_{0}\left(x_{1},-x_{2}\right), & \frac{\partial v_{R}^{1}}{\partial t}\left(x_{1}, x_{2}, 0\right)=-\tilde{v}_{1}\left(x_{1},-x_{2}\right), \\
v_{R}^{2}\left(x_{1}, x_{2}, 0\right)=-\tilde{v}_{0}\left(-x_{1}, x_{2}\right), & \frac{\partial v_{R}^{2}}{\partial t}\left(x_{1}, x_{2}, 0\right)=-\tilde{v}_{1}\left(-x_{1}, x_{2}\right), \\
v_{R}^{0}\left(x_{1}, x_{2}, 0\right)=\tilde{v}_{0}\left(-x_{1},-x_{2}\right), & \frac{\partial v_{R}^{0}}{\partial t}\left(x_{1}, x_{2}, 0\right)=\tilde{v}_{1}\left(-x_{1},-x_{2}\right) .
\end{array}
$$

Then the solution $u$ constructed in Theorem 3.2 can be equivalently decomposed:

$$
\begin{equation*}
u=u_{I}+u_{R}^{1}+u_{R}^{2}+u_{R}^{0}, \tag{3.26}
\end{equation*}
$$

where $u_{I}$ is the restriction to $\Omega$ of the solution $\tilde{u}_{I}$ of the wave equation without boundary, i.e., with ( $\tilde{u}_{0}, \tilde{u}_{1}$ being defined as were $\tilde{v}_{0}, \tilde{v}_{1}$ ),

$$
\begin{aligned}
& \frac{\partial^{2} \tilde{u}_{I}}{\partial t^{2}}-\Delta \tilde{u}_{I}=0, \\
& \tilde{u}_{I}(x, 0)=\tilde{u}_{0}(x), \\
& \frac{\partial \tilde{u}_{I}}{\partial t}(x, 0)=\tilde{u}_{1}(x),
\end{aligned}
$$

and where $u_{R}^{j}(j=1,2,0)$. denotes the restriction to $\Omega$ of the solution $\tilde{u}_{R}^{j}$ of

$$
\mathbf{L}\left(\tilde{u}_{R}^{j}\right)=\tilde{v}_{R}^{j}, \quad \tilde{u}_{R}^{j}(x, 0)=\frac{\partial \tilde{u}_{R}^{j}}{\partial t}(x, 0)=\frac{\partial^{2} \tilde{u}_{R}^{j}}{\partial t^{2}}(x, 0)=\frac{\partial^{3} \tilde{u}_{R}^{j}}{\partial t^{3}}(x, 0)=0 .
$$

The decomposition (3.26) can be interpreted as follows:

- $u_{I}$ is the incident wave.
- $u_{R}^{1}$ is the wave reflected by the absorbing boundary $\Gamma_{1}$.
- $u_{R}^{2}$ is the wave reflected by the absorbing boundary $\Gamma_{2}$.
- $u_{R}^{0}$ is the secondary reflected wave corresponding to the reflection of $u_{R}^{1}$ on $\Gamma_{2}$ and that of $u_{R}^{2}$ on $\Gamma_{1}$.

In the case of a point source, we can represent the different wave fronts corresponding to decomposition (3.26) as shown below. The three curves depicting $u_{I}$ indicate that the amplitude of $u_{I}$ is greater than that of $u_{R}^{1}$ and $u_{R}^{2}$, each represented by two curves, which is in turn greater than that of $u_{R}$ represented by a single curve. Also, along $u_{R}^{1}$ and $u_{R}^{2}$ the arrows point in the direction of decreasing amplitude (cf. numerical results of $\S 4$ ).

3.4. The case $\gamma \neq \frac{3}{2}$ : existence of a corner wave. We return now to the problem ( $\mathrm{P}_{\gamma}$ ) for arbitrary $\gamma>0$. We again assume that all initial data is $C^{\infty}$ and is compactly supported in $\Omega$. First, suppose that $u_{\gamma}$ is a solution to $\left(\mathrm{P}_{\gamma}\right)$ for some given $\gamma$. If we denote by $u^{*}$ the regular solution of $\mathrm{P}_{\gamma}$ for $\gamma=\frac{3}{2}$ we may define

$$
v_{\gamma}=u_{\gamma}-u^{*}
$$

Thus we can decompose $u_{\gamma}$ :

$$
u_{\gamma}=u^{*}+v_{\gamma},
$$

where $u^{*}$ is regular everywhere and $v_{\gamma}$ is a function that clearly satisfies the following equations (here we omit the index $\gamma$ for simplicity):

$$
\begin{array}{llll} 
& \frac{\partial^{2} v}{\partial t^{2}}-\Delta v=0, & x \in \Omega, & t>0, \\
& \frac{\partial^{2} v}{\partial t^{2}}+\frac{\partial^{2} v}{\partial t \partial x_{2}}-\frac{1}{2} \frac{\partial^{2} v}{\partial x_{1}^{2}}=0, & x \in \Gamma_{1}, \quad t>0 & \text { (CL1), } \\
& \frac{\partial^{2} v}{\partial t^{2}}+\frac{\partial^{2} v}{\partial t \partial x_{1}}-\frac{1}{2} \frac{\partial^{2} v}{\partial x_{2}^{2}}=0, & x \in \Gamma_{2}, \quad t>0 & (\mathrm{CL} 2), \\
& \gamma \frac{\partial^{2} v}{\partial t^{2}}+\frac{\partial}{\partial t}\left(\frac{\partial v}{\partial x_{1}}+\frac{\partial v}{\partial x_{2}}\right)=g, & x=0, \quad t>0 & (\mathrm{CC} \gamma)^{\prime}, \\
& v(x, 0)=\frac{\partial v}{\partial t}(x, 0)=0, & x \in \Omega,
\end{array}
$$

with

$$
\begin{equation*}
g=g_{\gamma} \equiv\left(\gamma-\gamma^{*}\right) \frac{\partial^{2} u^{*}}{\partial t^{2}} \tag{3.27}
\end{equation*}
$$

(We remark that, in this section, for simplicity of exposition we have chosen to work with the corner condition ( $\mathrm{CC} \gamma)^{\prime}$ instead of ( $\mathrm{CC} \gamma$ ), since with the zero initial conditions ( $\mathrm{CC} \gamma)^{\prime}$ and ( $\mathrm{CC} \gamma$ ) are equivalent.)

What we will do in the following is to show that there exists a solution $H_{\gamma}$ of $\tilde{P}_{\gamma}$ with $g=\delta(t)$. Then $v_{\gamma}$ will be defined to be $H_{\gamma^{*}} g_{\sigma}$ and will thus be a solution to ( $\tilde{\mathrm{P}}_{\gamma}$ ) with $g=g_{\gamma}$. We will show that $v_{\gamma}$ has a singularity at the corner and will say that $v_{\gamma}$ is the corner wave due to the corner condition ( $\mathrm{CC} \gamma)$. Finally, the solution $u_{\gamma}$ to $\left(\tilde{\mathrm{P}}_{\gamma}\right)$, with $g=0$ or equivalently to $\left(\mathrm{P}_{\gamma}\right)$, will be obtained by adding the regular solution $u^{*}$ of ( $\mathrm{P}_{\gamma^{*}}$ ) to $v_{\gamma}: u_{\gamma}=v_{\gamma}+u^{*}$.

To construct the solution $H_{\gamma}$ to $\left(\tilde{\mathrm{P}}_{\gamma}\right)$ we will proceed much as for the construction in $\S 3.3$ of the regular solution $u^{*}$ to $\left(\mathrm{P}_{\gamma^{*}}\right)$, except that here we need to introduce a singularity at the corner point. Thus we will start with the elementary solution $G$ to the wave equation in $\mathbb{R}^{2}$ at the point $(0,0)$. Then, as before, we will integrate $G$ along the characteristics of (3.19) to obtain $\tilde{G}$, which should satisfy the wave equation as well as the zero initial conditions of $\left(\tilde{\mathrm{P}}_{\gamma}\right)$. However, as $G$ does not vanish on the boundary $\Gamma$, there is no reason why $\tilde{G}$ should satisfy the boundary conditions (CL1) and (CL2) of ( $\tilde{\mathrm{P}}_{\gamma}$ ). However, we observe that $\partial^{2} G / \partial x_{1} \partial x_{2}$, which will play the same role here as $v$ in $\S 3.3$, does vanish on $\Gamma$. Therefore, taking the corresponding derivative of $\tilde{G}, H=\partial^{2} G / \partial x_{1} \partial x_{2}$, we will obtain a function $H$ that we will demonstrate satisfies all of the equations of $\left(\tilde{\mathrm{P}}_{\gamma}\right)$ except the corner condition. Moreover, we will show that a constant multiple $H_{\gamma}=a_{\gamma} H$ of $H$ does indeed satisfy the condition (CC $\left.\gamma\right)^{\prime}$.

To be more precise, we let $G$ denote the solution to

$$
\begin{array}{ll}
\frac{\partial^{2} G}{\partial t^{2}}-\Delta G=\delta(x) \times \delta(t), & x \in \mathbb{R}^{2}, \quad t>0, \\
G(x, 0)=\frac{\partial G}{\partial t}(x, 0)=0, & x \in \mathbb{R}^{2} .
\end{array}
$$

Then $G$ is given explicitly by

$$
G(x, t)= \begin{cases}\frac{1}{2} \pi\left(t^{2}-x_{1}^{2}-x_{2}^{2}\right)^{-1 / 2}, & x_{1}^{2}+x_{2}^{2}<t^{2} \\ 0 & \text { otherwise }\end{cases}
$$

Next we integrate $G$ twice in the direction $\left(t+x_{1}\right)$ and twice in the direction $\left(t+x_{2}\right)$ to obtain the solution $\tilde{G}$ to the problem

$$
\begin{array}{ll}
L \tilde{G}=\frac{1}{4}\left(\frac{\partial}{\partial t}+\frac{\partial}{\partial x_{1}}\right)^{2}\left(\frac{\partial}{\partial t}+\frac{\partial}{\partial x_{2}}\right)^{2} \tilde{G}=G, & x \in \mathbb{R}^{2}, \quad t>0, \\
\tilde{G}(x, 0)=\frac{\partial \tilde{G}}{\partial t}(x, 0)=\frac{\partial^{2} \tilde{G}}{\partial t^{2}}(x, 0)=\frac{\partial^{3} \tilde{G}}{\partial t^{3}}(x, 0)=0, & x \in \mathbb{R}^{2} . \tag{3.28}
\end{array}
$$

An explicit expression for $\tilde{G}$ may be obtained by straightforward calculation:

$$
\tilde{G}= \begin{cases}\frac{1}{12}\left(t-x_{1}-x_{2}\right)^{3} F(b(x, t)), & x_{1}^{2}+x_{2}^{2}<t^{2}  \tag{3.29}\\ 0 & \text { otherwise }\end{cases}
$$

where

$$
\begin{align*}
& F(b)=\int_{0}^{b} \int_{0}^{s} \frac{\sigma^{3 / 2}}{(1+\sigma)^{2}} d \sigma d s, \\
& b(x, t)=\frac{t^{2}-x_{1}^{2}-x_{2}^{2}}{\left(t-x_{1}-x_{2}\right)^{2}} . \tag{3.30}
\end{align*}
$$

Remark. We may further calculate that

$$
\begin{equation*}
F(b)=\left(5+\frac{4}{3} b\right) \sqrt{b}-(5+3 b) \arctan \sqrt{b} \tag{3.31}
\end{equation*}
$$

with $b$ given by (3.30); however, as we are interested in the derivatives of $\tilde{G}$, the form (3.30) is easier to work with.

Concerning the regularity of $\tilde{G}$, we note that $\tilde{G}$ is bounded but singular along the plane $x_{1}+x_{2}=t$. However, the intersection of this plane with the domain that interests us, $\bar{\Omega} \times \mathbb{R}^{+}$, is the point $x=0, t=0$.

Next we introduce the function $H$ defined by

$$
H(x, t)=\frac{\partial^{2} \tilde{G}}{\partial x_{1} \partial x_{2}}(x, t) .
$$

We can show the following result.
Theorem 3.4. (i) $H$ vanishes in the part of $\bar{\Omega} \times \mathbb{R}^{+}$outside the closure of the cone $x_{1}^{2}+x_{2}^{2}<t^{2}$.
(ii) $H$ is $C^{\infty}$ in the part of $\bar{\Omega} \times \mathbb{R}^{+}$.
(iii) $H$ is $C^{1}$ in $\bar{\Omega} \times \mathbb{R}^{+} \backslash\{(0,0,0)\}$.
(iv) The second derivatives of $H$ are singular along $x_{1}^{2}+x_{2}^{2}=t^{2}$.
(v) $H$ satisfies the wave equation in the distributional sense in $\mathbb{R}^{2} \times \mathbb{R}_{*}^{+}$and in the classical sense in $\bar{\Omega} \times \mathbb{R}_{*}^{+}$except along $x_{1}^{2}+x_{2}^{2}=t^{2}$. In particular

$$
\frac{\partial^{2} H}{\partial t^{2}}-\Delta H=0, \quad(x, t) \in \bar{\Omega} \times \mathbb{R}^{+} \cap\left\{(x, t): x_{1}^{2}+x_{2}^{2}<t\right\} .
$$

$H$ also satisfies the initial conditions

$$
H(x, 0)=\frac{\partial H}{\partial t}(x, 0)=0, \quad x \in \Omega .
$$

(vi) H satisfies the boundary conditions (CL1) and (CL2).
(vii) $H$ satisfies $\left(\mathrm{CC}_{\gamma^{*}}\right)^{\prime}$ with $g=0$ for $t>0$.

Proof. See Theorem 3.3, Corollary 3.1, and Lemmas 3.2 and 3.3 of [1] for the details. The idea is the same as that for the smooth solution given in Theorem 3.2, the difference being that here we have introduced a singularity via the elementary solution $G$.

We point out that while $H$ satisfies $\left(\mathrm{CC}_{\gamma^{*}}\right)^{\prime}$ with $g=0$ for $t>0$, it does not satisfy $\left(\mathrm{CC}_{\gamma^{*}}\right)$. In fact we establish the following important result.

Lemma 3.3. The first derivatives of $H$ in time and in space at the corner $x_{1}=x_{2}=0$, for $t>0$, are constant in time and positive, i.e.,

$$
\begin{aligned}
& \frac{\partial H}{\partial t}(0, t)=c_{1}>0 \quad \text { for } t>0, \\
& \frac{\partial H}{\partial x_{1}}(0, t)=\frac{\partial H}{\partial x_{2}}(0, t)=c_{2}>0 \quad \text { for } t>0 .
\end{aligned}
$$

Thus for $\gamma>0$, we have along the axis $(0,0, t), t>0$,

$$
\frac{d}{d t}\left(\gamma \frac{\partial H}{\partial t}+\frac{\partial H}{\partial x_{1}}+\frac{\partial H}{\partial x_{2}}\right)=\frac{1}{a_{\gamma}} \delta(0),
$$

where the amplitude $a_{\gamma}$ is defined by

$$
\frac{1}{a_{\gamma}}=\gamma c_{1}+2 c_{2}>0 .
$$

Proof. We consider first the time derivative. From (3.29), (3.30) we calculate

$$
\begin{align*}
& H=\frac{\partial^{2}}{\partial x_{1} \partial x_{2}} \tilde{G}=\frac{1}{2}\left(t-x_{1}-x_{2}\right) F(b)-\frac{1}{4}\left(t-x_{1}-x_{2}\right)^{2} F^{\prime}(b)\left(\frac{\partial b}{\partial x_{1}}+\frac{\partial b}{\partial x_{2}}\right) \\
&+\frac{1}{12}\left(t-x_{1}-x_{2}\right)^{3}\left(F^{\prime \prime}(b) \frac{\partial b}{\partial x_{1}} \frac{\partial b}{\partial x_{2}}+F^{\prime}(b) \frac{\partial^{2} b}{\partial x_{1} \partial x_{2}}\right) . \tag{3.32}
\end{align*}
$$

Now for $x_{1}=x_{2}=0$ and $t>0, b=1$ while $\partial b / \partial x_{1}=\partial b / \partial x_{2}=t / 2$ and $\partial^{2} b / \partial x_{1} \partial x_{2}=6 / t^{2}$. Thus

$$
H(0, t)=\frac{t}{4}\left(2 F(1)-2 F^{\prime}(1)+\frac{4}{3} F^{\prime \prime}(1)\right), \quad t>0
$$

and we conclude that $\partial H / \partial t(0, t)$ is constant, $t>0$. More precisely, we have

$$
F(1)-F^{\prime}(1)=-\int_{0}^{1} \frac{\sigma^{5 / 2}}{(1+\sigma)^{2}} d \sigma=\frac{23}{6}-\frac{5 \pi}{4}, \quad F^{\prime \prime}(1)=\frac{1}{4},
$$

from which we deduce that

$$
\frac{\partial H}{\partial t}(0, t)=2-\frac{5 \pi}{8}>0 .
$$

As $H$ is symmetric in the variables $x_{1}, x_{2}$, we clearly have $\partial H / \partial x_{1}=\partial H / \partial x_{2}$ at $x=0$. We can obtain an expression for $\partial H / \partial x_{1}$ from (3.29) by straightforward calculation that permits us to determine that $\partial H / \partial x_{1}$ is constant in time at $x_{1}=x_{2}=0, t>0$, but that is not very amenable to determining that the constant is positive. Thus it is simpler to proceed in an indirect fashion.

As $H$ satisfies the wave equation in $\Omega, t>0$, in the distributional sense, on integrating over $\Omega$ and differentiating in time we obtain

$$
\frac{d^{3}}{d t^{3}} \int_{\Omega} H d x_{1} d x_{2}-\frac{d}{d t} \int_{\Gamma} \frac{\partial H}{\partial \nu} d \gamma=0, \quad t>0 .
$$

Integrating conditions (CL1) over $\Gamma_{1}$ and (CL2) over $\Gamma_{2}$ and adding, we have

$$
\frac{d^{2}}{d t^{2}} \int_{\Gamma} H d \gamma+\frac{d}{d t} \int_{\Gamma} \frac{\partial H}{\partial \nu} d \gamma=\frac{1}{2}\left[\frac{\partial H}{\partial x_{1}}(0,0, t)+\frac{\partial H}{\partial x_{2}}(0,0, t)\right] .
$$

We then combine these two equations to arrive at

$$
\frac{d^{3}}{d t^{3}} \int_{\Omega} H d x_{1} d x_{2}+\frac{d^{2}}{d t^{2}} \int_{\Gamma} H d \gamma=\frac{1}{2}\left[\frac{\partial H}{\partial x_{1}}(0,0, t)+\frac{\partial H}{\partial x_{2}}(0,0, t)\right] .
$$

The relation between $H$ and $\tilde{G}, H=\partial^{2} \tilde{G} / \partial x_{1} \partial x_{2}$, gives us

$$
\frac{\partial^{3}}{\partial t^{3}} \tilde{G}(0,0, t)+\frac{\partial^{2}}{\partial t^{2}}\left[\frac{\partial \tilde{G}}{\partial x_{1}}(0,0, t)+\frac{\partial \tilde{G}}{\partial x_{2}}(0,0, t)\right]=\frac{1}{2}\left[\frac{\partial H}{\partial x_{1}}(0,0, t)+\frac{\partial H}{\partial x_{2}}(0,0, t)\right] .
$$

We calculate from (3.28), (3.29)

$$
\begin{aligned}
& \tilde{G}(0,0, t)=\frac{1}{12} t^{3} F(1) \\
& \frac{\partial \tilde{G}}{\partial x_{1}}(0,0, t)=\frac{\partial \tilde{G}}{\partial x_{2}}(0,0, t)=\frac{t^{2}}{4}\left[-F(1)+\frac{2}{3} F^{\prime}(1)\right]
\end{aligned}
$$

Thus we have

$$
\begin{aligned}
\frac{\partial H}{\partial x_{1}}(0,0, t) & =\frac{\partial H}{\partial x_{2}}(0,0, t) \\
& =-\frac{1}{2} F(1)+\frac{2}{3} F^{\prime}(1) \\
& =\frac{1}{2} \int_{0}^{1} \frac{\left(\sigma+\frac{1}{3}\right) \sigma^{3 / 2}}{(1+\sigma)^{2}} d \sigma \\
& =\frac{\pi-3}{2}>0
\end{aligned}
$$

It is now natural to define

$$
\begin{equation*}
H_{\gamma}=a_{\gamma} H \tag{3.33}
\end{equation*}
$$

and it follows that

$$
\frac{d}{d t}\left(\gamma \frac{\partial H_{\gamma}}{\partial t}+\frac{\partial H_{\gamma}}{\partial x_{1}}+\frac{\partial H_{\gamma}}{\partial x_{2}}\right)=\delta(0)
$$

along the half-line $(0,0, \mathrm{t}), t>0$. Thus $H_{\gamma}$ is a solution to $\tilde{P}_{\gamma}$ with $g=\delta(t)$. We next put

$$
\begin{equation*}
v_{\gamma}=H_{\gamma} * g_{\gamma} \tag{3.34}
\end{equation*}
$$

where $g_{\gamma}$ is given by (3.27), and finally define

$$
\begin{equation*}
u_{\gamma}=v_{\gamma}+u^{*} \tag{3.35}
\end{equation*}
$$

THEOREM 3.5. The function $u_{\gamma}$ given by (3.35) is the unique weak solution of $P_{\gamma}$. $u^{*}$ is the regular part of the solution and $v_{\gamma}$ is the singular corner wave, whose properties are:
(i) $v_{\gamma}$ is of finite energy for any $t>0$;
(ii) $v_{\gamma}$ is of class $C^{1}$ in $\bar{\Omega} \times \mathbb{R}^{+}$;
(iii) $v_{\gamma}$ has singular second-order spatial derivatives along (a) the line $x=0$, and (b) the cone $|x|=t$.

Proof. See [1] for the proof.
We conclude this section with three important remarks.
Remark 3.1. For each positive $\gamma, u_{\gamma}$ is proportional to $u^{*}$ along the axis $x_{1}=x_{2}=0$. Thus $u_{\gamma}$ is $C^{\infty}$ in time along this axis.

Remark 3.2. For each positive $\gamma, u_{\gamma}$ is a solution of (1.1), (1.5), and (1.7) in the class of finite energy functions. We thus have a counterexample to the uniqueness of the solution of the problem without a corner condition.

Remark 3.3. It is interesting to note that for each positive $\gamma$, the corner wave $v_{\gamma}$ is proportional to $H * \partial^{2} u^{*} / \partial t^{2}$ :

$$
v_{\gamma}=\frac{\gamma-\gamma^{*}}{\gamma c_{1}+2 c_{2}} H * \frac{\partial^{2} u^{*}}{\partial t^{2}}
$$

and that the constant of proportionality remains bounded between $-\left(\gamma^{*} / 2 c_{2}\right)$ and $1 / c_{1}$ as $\gamma$ ranges between 0 and $+\infty$, the constant being negative for $\gamma<\gamma^{*}$, positive for $\gamma>\gamma^{*}$, and vanishing for $\gamma=\gamma^{*}$. In fact we see that $v_{\gamma}$ is well defined even for $\gamma \leqq 0$ except for $\gamma=-2\left(c_{2} / c_{1}\right)$.

## 4. Numerical results.

4.1. Numerical scheme. Here we exhibit the results of numerical experiments confirming the conclusions obtained in the previous section. These results were obtained using a variational scheme, which we apply not to the wave equation (1.1) itself but to its time derivative so we can use the boundary equations after the standard integration by parts in the variational formulation. The discretization in time is obtained using an explicit scheme incorporating the four time levels $n-2, n-1, n$, and $n+1$, as a third-order time derivative appears in the time-differentiated wave equation. We remark that it is not the numerical scheme per se that interests us here, and that another scheme-the finite difference scheme proposed by Engquist and Majda in [5], for example-could just as well have been used.

We let $\Omega$ denote the unit square $(0,1) \times(0,1)$ with boundary $\Gamma$ written as the disjoint union of $\Gamma_{1}$ the upper edge, $\Gamma_{2}$ the right-hand edge, $C$ the upper right-hand corner, and $\Gamma_{R}$ the left and lower edges with the three remaining corners (see Fig. 1).


Fig. 1. Domain $\Omega$ and boundary $\Gamma=\Gamma_{1} \cup \Gamma_{2} \cup C \cup \Gamma_{R}$.
The problem we will solve numerically is the following:

$$
\begin{equation*}
\frac{\partial^{3} u}{\partial t^{3}}-\frac{\partial \Delta u}{\partial t}=\frac{\partial f}{\partial t} \quad \text { in } \Omega, \tag{4.1}
\end{equation*}
$$

with boundary condition

$$
\begin{equation*}
\frac{\partial^{2} u}{\partial t^{2}}+\frac{\partial^{2} u}{\partial t \partial x_{2}}-\frac{1}{2} \frac{\partial^{2} u}{\partial x_{1}^{2}}=0 \quad \text { in } \Gamma_{1}, \tag{4.2}
\end{equation*}
$$

$$
\begin{gather*}
\frac{\partial^{2} u}{\partial t^{2}}+\frac{\partial^{2} u}{\partial t \partial x_{1}}-\frac{1}{2} \frac{\partial^{2} u}{\partial x_{2}^{2}}=0 \quad \text { in } \Gamma_{2},  \tag{4.2}\\
\gamma \frac{\partial u}{\partial t}+\frac{\partial u}{\partial x_{1}}+\frac{\partial u}{\partial x_{2}}=0 \quad \text { at } C,  \tag{4.3}\\
u=0 \quad \text { on } \Gamma_{R}, \tag{4.4}
\end{gather*}
$$

and initial conditions

$$
\begin{array}{cc}
u=u_{0} & \text { at } t=0,  \tag{4.5}\\
\frac{\partial u}{\partial t}=u_{1} & \text { at } t=0,
\end{array}
$$

where $f \in L^{2}(\Omega)$ is of compact support.
The solution $u$ will be sought in the space $W$ of functions in $H^{1}(\Omega)$ having trace in $H_{0}^{1}(\Gamma)$, where by $H_{0}^{1}(\Gamma)$ we will denote the Hilbert space of functions in $L^{2}(\Gamma)$ whose restrictions to $\Gamma_{1}$ and to $\Gamma_{2}$ are $H^{1}$ functions determining the same value at the corner $C$ and whose restriction to $\Gamma_{R}$ is identically zero (see § 3.1):

$$
\begin{aligned}
& W=\left\{v \in H^{1}(\Omega): \operatorname{trace}(u) \in H_{0}^{1}(\Gamma)\right\}, \\
& H_{0}^{1}(\Gamma)=\left\{\phi \in L^{2}(\Gamma): \phi_{1}=\phi_{\mid \Gamma_{1}} \in H^{1}\left(\Gamma_{1}\right) ; \phi_{2}=\phi_{\mid \Gamma_{2}} \in H^{1}\left(\Gamma_{2}\right) ;\right. \\
& \left.\qquad \phi_{1}(C)=\phi_{2}(C) ; \text { and } \phi_{\mid \Gamma_{R}} \equiv 0\right\} .
\end{aligned}
$$

Thus, to obtain a variational form of (4.1), we multiply by a test function $v \in W$ and integrate over $\Omega$ :

$$
\left(\frac{\partial^{3} u}{\partial t^{3}}-\frac{\partial \Delta u}{\partial t}, v\right)=\left(\frac{\partial f}{\partial t}, v\right) \quad \text { for all } v \in W \text {, }
$$

where (, ) denotes the inner product in $L^{2}(\Omega)$.
Integrating by parts, we have

$$
\begin{equation*}
\frac{\partial^{3}}{\partial t^{3}}(u, v)+\frac{\partial}{\partial t}(\nabla u, \nabla v)-\left\langle\frac{\partial^{2} u}{\partial t \partial \nu}, v\right\rangle_{\Gamma}=\frac{\partial}{\partial t}(f, v) \quad \text { for all } v \in V, \tag{4.6}
\end{equation*}
$$

where $\langle,\rangle_{\Gamma}$ denotes the inner product in $L^{2}(\Gamma)$ and $\nu$ is the outward pointing unit normal vector on $\Gamma$. The boundary conditions (4.2) and (4.4) imply

$$
\begin{equation*}
-\left\langle\frac{\partial^{2} u}{\partial t \partial \nu}, v\right\rangle_{\Gamma}=\frac{\partial^{2}}{\partial t^{2}}\langle u, v\rangle_{\Gamma}-\frac{1}{2}\left\langle\frac{\partial^{2} u}{\partial x_{1}^{2}}, v\right\rangle_{\Gamma_{1}}-\frac{1}{2}\left\langle\frac{\partial^{2} u}{\partial x_{2}^{2}}, v\right\rangle_{\Gamma_{2}}, \tag{4.7}
\end{equation*}
$$

where $\langle,\rangle_{\Gamma_{1}}$ and $\langle,\rangle_{\Gamma_{2}}$ have the obvious meanings, and integration by parts yields

$$
\begin{gather*}
-\left\langle\frac{\partial^{2} u}{\partial t \partial \nu}, v\right\rangle_{\Gamma}=\frac{\partial^{2}}{\partial t^{2}}\langle u, v\rangle_{\Gamma}+\frac{1}{2}\left\langle\frac{\partial u}{\partial x_{1}}, \frac{\partial v}{\partial x_{1}}\right\rangle_{\Gamma_{1}}+\frac{1}{2}\left\langle\frac{\partial u}{\partial x_{2}}, \frac{\partial v}{\partial x_{2}}\right\rangle_{\Gamma_{2}}  \tag{4.8}\\
-\frac{1}{2}\left(\frac{\partial u}{\partial x_{1}}(1,1)+\frac{\partial u}{\partial x_{2}}(1,1)\right) v(1,1) .
\end{gather*}
$$

Plugging (4.8) into (4.6) and applying the corner condition (4.3), we arrive at the
variational problem on which our numerical scheme is based: Find $u \in W$ such that

$$
\begin{align*}
\frac{\partial^{3}}{\partial t^{3}}(u, v) & +\frac{\partial}{\partial t}(\nabla u, \nabla v)+\frac{\partial^{2}}{\partial t^{2}}\langle u, v\rangle_{\Gamma}+\frac{1}{2}\left\langle\frac{\partial u}{\partial x_{1}}, \frac{\partial v}{\partial x_{1}}\right\rangle_{\Gamma_{1}}+\frac{1}{2}\left\langle\frac{\partial u}{\partial x_{2}}, \frac{\partial v}{\partial x_{2}}\right\rangle_{\Gamma_{2}}  \tag{4.9}\\
& +\frac{1}{2} \gamma \frac{\partial}{\partial t} u(1,1) v(1,1)=\frac{\partial}{\partial t}(f, v) \quad \forall v \in W .
\end{align*}
$$

To discretize $\Omega$ we use a uniform mesh of triangles obtained by cutting diagonally, from the upper left to the lower right, each of the squares of a uniform mesh of squares.

The approximation $u_{h}$ of $u$ will be sought in the space $W_{h}$ of functions in $W$ whose restriction to each triangle of the mesh is linear. Thus the degrees of freedom are the values at the vertices. Mass lumping is used for the integrals not involving spatial derivatives. Other integrals are computed exactly.

The time discretization is an explicit four-level scheme involving times $m, m+1$, $m-1$, and $m-2$ and is centered at time $m-\frac{1}{2}$. For more details see [1].
4.2. Description of experiments. In these experiments we are concerned with the reflection by the corner. We study the corner condition (4.3) for several choices of the constant $\gamma, \gamma=1.5, \gamma=0.1$, and $\gamma=3.0$.

To generate strong reflections at the corner, the source was placed near the upper right-hand corner $C=(1,1)$ of the domain $\Omega=[0,1] \times[0,1]$ with center at $(.89, .89)$ and radius .04 (see Fig. 2). The source was introduced as a right-hand side, $f\left(x_{1}, x_{2}, t\right)=$ $g\left(x_{1}, x_{2}\right) h(t)$ :

$$
g\left(x_{1}, x_{2}\right)= \begin{cases}10,000 \times(1-r / .04) & \text { if } r<.04 \\ 0 & \text { otherwise }\end{cases}
$$



Fig. 2. Source and location of seismograms.
where $r$ is the distance of the point $\left(x_{1}, x_{2}\right)$ from the center of the source $(.89, .89)$,

$$
h(t)= \begin{cases}e^{-10}(1-t / .05)^{2} & \text { if } t<.1 \\ 0 & \text { otherwise }\end{cases}
$$

The calculations were done with meshsize $\Delta x_{1}=\Delta x_{2}=.01$ and timestep $\Delta t=.005$.
The results of the experiments are presented first in the form of "snapshots" representing the displacement $u$ as a function of $\left(x_{1}, x_{2}\right)$ at the times $t=.3$ and $t=.5$, and then as seismograms representing the displacement $u$ as a function of time $t, 0 \leqq t \leqq 1$, at the points $\mathrm{P} 1=(.79, .79) \mathrm{P} 2=(.99, .79)$, and $\mathrm{P} 3=(.99, .99)$ (see Fig. 2).

The solutions obtained with the various choices of $\gamma$ are compared with the so-called "exact" solution, obtained by doing the calculations on the larger rectangle $[0,2] \times[0,2]$ so that the wave has not reached any edge and thus there is no reflection by the observation times $t=.3$ and $t=.5$, and so that the reflection has not reached any of the observation points $\mathrm{P} 1, \mathrm{P} 2$, or P 3 by the end time of the experiment, $t=1$. We have also presented as seismograms results obtained by using the first-order absorbing boundary condition for the boundaries $\Gamma_{1}$ and $\Gamma_{2}$. To obtain these results, (4.2) ${ }_{1,2}$ are replaced by

$$
\begin{array}{ll}
\frac{\partial^{2} u}{\partial t^{2}}+\frac{\partial^{2} u}{\partial t \partial x_{2}}=0 & \text { on } \Gamma_{1}, \\
\frac{\partial^{2} u}{\partial t^{2}}+\frac{\partial^{2} u}{\partial t \partial x_{1}}=0 & \text { on } \Gamma_{2},
\end{array}
$$

so that (4.7) becomes

$$
-\left\langle\frac{\partial^{2} u}{\partial t \partial \nu}, v\right\rangle_{\Gamma}=\frac{\partial^{2}}{\partial t^{2}}\langle u, v\rangle_{\Gamma},
$$

and (4.9) is replaced by

$$
\frac{\partial^{3}}{\partial t^{3}}(u, v)+\frac{\partial}{\partial t}(\nabla u, \nabla v)+\frac{\partial^{2}}{\partial t^{2}}\langle u, v\rangle_{\Gamma}=\frac{\partial}{\partial t}(f, v) \quad \text { for all } v \in W \text {. }
$$

4.3. Experimental results. In Figs. 3-10 we see snapshots representing the displacement $u$ as a function of $\left(x_{1}, x_{2}\right)$ at times $t=.3$ and $t=.5$. Figures 3 and 4 give the exact solution. In Figs. 5-10, to see the reflected wave better, we have pictured the difference between the solution obtained with the second-order absorbing boundary condition for the sides $x_{1}=1$ and $x_{2}=1$ and the corner condition for the various choices of $\gamma$, and the exact solution.

In Figs. 5, 7, and 9, for $t=.3$ we see at a distance of about $\frac{1}{3}$ from the corner $C$ near both boundaries $x_{1}=1$ and $x_{2}=1$, a reflection that is due to the second-order absorbing boundary condition. The amplitude of this reflection is .14 , or $16 \%$ of the amplitude .81 of the initial wave at time $t=.3$. In Figs. 7 and 9 there appears much nearer the corner a reflection due to the corner that is the $\gamma$-wave. In Fig. 7, for $\gamma=0.1$, $\gamma$ too small, the $\gamma$-wave is positive of amplitude .25 , or 31 percent of the amplitude of the initial wave, whereas in Fig. 9, for $\gamma=3.0, \gamma$ too large, the $\gamma$-wave is negative of amplitude .15 , or 19 percent of the amplitude of the original wave.

In Figs. 6, 8, and 10, for $t=5$ we see the reflection due to the second-order boundary condition now at a distance about $\frac{1}{2}$ from the corner. The amplitude of the reflection is now .23 , or 38 percent of the amplitude .61 of the initial wave at time $t=.5$. (This increase in the amplitude of the reflection is expected, since at time $t=.5$


Fig. 3. Exact solution. Time $t=.3$; minimum $=0.00$; maximum $=0.81$.


Fig. 4. Exact solution. Time $t=.5 ;$ minimum $=0.00$; maximum $=0.61$.

Fig. 5. Difference between solution calculated with $\gamma=1.5$ and exact solution. Time $t=.3$; minimum $=$ -0.06 ; maximum $=0.08$.


Fig. 6. Difference between solution calculated with $\gamma=1.5$ and exact solution. Time $t=.5$; minimum $=$ -0.13 ; maximum $=0.10$.


Fig. 7. Difference between solution calculated with $\gamma=0.1$ and exact solution. Time $t=.3$; minimum $=$ -0.05 ; maximum $=0.24$.


Fig. 8. Difference between solution calculated with $\gamma=0.1$ and exact solution. Time $t=.5 ;$ minimum $=0.13$; maximum $=0.10$.

Fig. 9. Difference between solution calculated with $\gamma=3.0$ and exact solution. Time $t=.3$; minimum $=$ -0.18 ; maximum $=0.08$.


Fig. 10. Difference between solution calculated with $\gamma=3.0$ and exact solution. Time $t=.5$; minimum $=$ -0.15 ; maximum $=0.10$.
the incident wave is striking the boundary at an angle from the normal much larger than at time $t=.3$.) In Fig. 8 the $\gamma$-wave for small $\gamma$ is now of amplitude .15 , or 25 percent of the amplitude of the incident wave. In Fig. 10 the $\gamma$-wave for large $\gamma$ is of amplitude .10 , or 16 percent of the amplitude of the incident wave.

Thus at the earlier time we see that the reflection due to the corner, when $\gamma$ is improperly chosen, is as large as or larger than that due to the second-order condition itself. At the later time, however, it has diminished, whereas the reflection due to the second-order condition has increased. It is important to keep in mind, though, that the large amplitude reflection due to the second-order boundary condition is caused by waves striking the boundary at large angles from the normal. Thus the large reflection propagates at angles close to the tangential and does not travel very rapidly into the interior of the domain, while the reflection due to the corner travels directly toward the center of the domain.

We remark that in Figs. 5 and 6, where the "good" $\gamma, \gamma=1.5$, has been used, there seems also to be a slight reflection from the corner, though it is much smaller than the $\gamma$-waves appearing in Figs. 7-10. This is, however, only the effect of the second-order absorbing boundary condition.

The last six figures, Figs. 11-16 ${ }^{1}$, are the seismograms giving the response at three points at equal distances from the source, $\mathrm{P} 1=(.79, .79)$ in the interior, $\mathrm{P} 2=(.99, .79)$


Fig. 11. Response at $P_{1}=(.79, .79)$. Exact solution; second-order $\mathrm{ABC}(\gamma=1.5)$; first-order ABC .
near the boundary, and $\mathrm{P} 3=(.99, .99)$ near the corner. For the odd-numbered figures the solid line represents the exact solution, the dotted line the solution obtained with the second-order condition with $\gamma=1.5$, and the dashed line the solution obtained with the first-order condition. At all of the points, even at P1 which is farthest from the boundary, the better behavior of the second-order condition is quite pronounced.

For the even-numbered figures the three curves all represent solutions obtained with the second-order condition but with different choices of the constant $\gamma$. For the solution represented by the solid line we have $\gamma=1.5$, by the dotted line $\gamma=0.1$, and

[^1]

Fig. 12. Response at $P_{1}=(.79, .79)$. Second-order $\mathrm{ABC} \gamma=1.5$; second-order $\mathrm{ABC} \gamma=0.1$; second-order $\mathrm{ABC} \gamma=3.0$.


Fig. 13. Response at $P_{2}=(.99, .79)$. Exact solution; second-order $\mathrm{ABC}(\gamma=1.5)$; first-order ABC .


Fig. 14. Response at $P_{2}=(.99, .79)$. Second-order $\mathrm{ABC} \gamma=1.5$; second-order $\mathrm{ABC} \gamma=0.1$; second-order $\mathrm{ABC} \gamma=3.0$.


Fig. 15. Response at $P_{3}=(.99, .99)$. Exact solution; second-order $\mathrm{ABC}(\gamma=1.5)$; first-order ABC .


Fig. 16. Response at $P_{3}=(.99, .99)$. Second-order $\mathrm{ABC} \gamma=1.5$; second-order $\mathrm{ABC} \gamma=0.1$; second-order $\mathrm{ABC} \gamma=3.0$.
by the dashed line $\gamma=3.0$. Again at each point it is quite clear, especially on comparison with the exact solution in the corresponding odd-numbered figure, that the better solution is obtained with $\gamma=1.5$. We do remark that by the final time of the observation all of the solutions with the second-order condition have nearly converged to the exact solution, while the solution with the first-order condition remains at some distance.

As a final remark, we point out that we have not presented results with the value $\gamma=\sqrt{2}$ as proposed by Engquist and Majda [5], since with the scale we used we could hardly see the difference between results obtained with $\gamma=\sqrt{2}$ and those obtained with $\gamma=1.5$. In fact, for the two-dimensional case with the second-order absorbing boundary condition, the solution we have presented differs little from that proposed in [5], though it was obtained from a very different point of view. We have seen that the second-order condition is much better than the first-order condition even in the presence of a corner, and that the condition chosen for the corner is significant. Also we have
shown theoretically that our approach to the corner problem yields a solution to the problem in three dimensions and can be used to obtain a solution for the corner problem when higher-order boundary conditions are used.

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[^1]:    ${ }^{1}$ In Figs. 11-16, "absorbing boundary condition" is abbreviated "ABC."

